

Lifshitz holography for undoped Weyl semimetalsUmut Gürsoy[†], Vivian Jacobs^{*}, Erik Plauschinn^{*‡}, Henk Stoof^{*}, Stefan Vandoren^{*}[†] *Theory Group, Physics Department, CERN, CH-1211 Geneva 23, Switzerland*^{*} *Institute for Theoretical Physics and Spinoza Institute, Utrecht University
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Abstract

We continue our recently proposed holographic description of single-particle correlation functions for four-dimensional chiral fermions with Lifshitz scaling at zero chemical potential, paying particular attention to the dynamical exponent $z = 2$. We present new results for the spectral densities and dispersion relations at non-zero momenta and temperature. In contrast to the relativistic case with $z = 1$, we find the existence of a quantum phase transition from a non-Fermi liquid into a Fermi liquid in which two Fermi surfaces spontaneously form, even at zero chemical potential. Our findings show that the boundary system behaves much like an undoped Weyl semimetal.

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Contents

1	Introduction	3
2	Single-particle Green's function for Lifshitz fermions	5
3	Single-particle spectra	10
3.1	Relativistic case $z = 1$	10
3.1.1	Zero temperature	10
3.1.2	Non-zero temperature	11
3.1.3	Dispersion relation	13
3.2	Non-relativistic case $z \neq 1$	14
3.2.1	Zero temperature	14
3.2.2	Non-zero temperature	15
3.2.3	Dispersion relation	16
4	Quantum phase transition	20
4.1	Momentum distribution	20
4.2	Fermi momentum	21
4.2.1	Quasi-particle residue	21
4.2.2	Effective mass	23
4.2.3	Quasi-particle decay rate	26
4.3	Second-order WKB analysis around the Fermi surface	28
4.4	Phase diagram	31
4.5	Fermi surfaces and the quantum phase transition for $\frac{1}{2} < M < \frac{z}{2}$	34
5	Conclusion and discussion	36
A	Conventions	38
A.1	A note on dimensions and units	38
A.2	Vielbeins and Dirac matrices	38
B	Details on the near-boundary asymptotics	40
B.1	Dirac equation	40
B.2	Asymptotic solutions near the boundary	41
B.3	Hydrodynamic limit	43
C	Symmetry properties of the Green's function	43
C.1	Scale-invariant variables	43
C.2	Chirality	45
C.3	Particle-hole symmetry	45
C.4	A symmetry relating M to $-M$	45
C.5	A symmetry of the momentum distribution	46

D Pole structure and sum rules	47
D.1 Pole structure	47
D.2 Sum rule for $\vec{k} = 0$	48
D.3 Sum rule for $\vec{k} \neq 0$	49
E WKB calculation	50

1 Introduction

Over the past years, the AdS/CFT correspondence has become a more and more popular and widespread tool which offers the opportunity to apply ideas from string theory to realistic materials studied in condensed-matter physics, see e.g. [1, 2, 3, 4, 5] and references therein. More specifically, it is potentially very useful when considering strongly coupled or critical condensed matter, which is generically not possible to describe within perturbation theory. However, in the context of the AdS/CFT correspondence these systems can be understood and investigated more conveniently using a dual theory in a higher-dimensional curved spacetime, i.e., in the framework of general relativity. Most commonly, the duality exists between a weakly coupled, classical gravity theory in a curved anti-de-Sitter (AdS) spacetime on the one hand, and a strongly coupled, conformal field theory (CFT) living on the flat boundary of the AdS spacetime on the other hand. Such a conformal field theory describes for example a quantum critical point of the condensed-matter system under consideration, and the observables that are most readily available from the AdS/CFT correspondence are usually correlation functions of the *composite* operators that classify the conformal field theory. For our purposes, most relevant are correlators of fermionic operators, since they may show Fermi or non-Fermi liquid-like behavior most easily [6, 7, 8, 9, 10].

However, in condensed-matter physics it is more natural to think in terms of fermionic *single-particle* operators, i.e., creation and annihilation operators in a Fock space, that satisfy the following (equal-time) anti-commutation relations

$$[\psi_\alpha(\vec{x}, t), \psi_{\alpha'}^\dagger(\vec{x}', t)]_+ = \delta(\vec{x} - \vec{x}')\delta_{\alpha, \alpha'} , \quad (1.1)$$

where α labels the spin of the electron. Furthermore, in experiments on condensed-matter systems, the quantity that is measured is essentially always related to *single-particle* or *two-particle* correlation functions and not to the above-mentioned composite operators. For example, the retarded single-particle correlation or Green's function G_R is measurable in electronic systems using angle-resolved photoemission spectroscopy (ARPES). For this reason, we are interested in finding G_R of a strongly interacting condensed-matter system using holographic methods. Inspired by previous work on this topic, such as [11] and [12, 13], we have shown recently how the usual holographic prescription can be modified in such a manner that it allows for the construction of the retarded *single-particle* Green's function [14]. In a sense, our method can be seen as a bulk derivation of the semi-holographic description advocated in [12].

A crucial consequence of the anti-commutation relations in equation (1.1) is the existence of a sum rule for the corresponding single-particle two-point correlation function, given by

$$\frac{1}{\pi} \int_{-\infty}^{+\infty} d\omega \operatorname{Im} \left[G_{R;\alpha,\alpha'}(\vec{k}, \omega) \right] = \delta_{\alpha,\alpha'} . \quad (1.2)$$

This sum rule is essential in determining whether the quantity under consideration is the correlator of a single-particle or a composite operator, and therefore plays a central role in our construction.

The quantum critical points described by the AdS/CFT correspondence are characterized by their dynamical scaling exponent z . In the usual correspondence, an anti-de-Sitter background leads to a quantum critical point with relativistic scaling, i.e., $z = 1$. But from a condensed-matter perspective we are also interested in quantum critical points that have a dynamical scaling exponent different from one. To achieve this, the usual anti-de-Sitter background is generalized into a so-called Lifshitz background, which leads to a non-relativistic, i.e., $z \neq 1$ scaling on the boundary [15]. Non-zero temperature effects can then be studied by placing a black brane in the Lifshitz spacetime [16], see also [17] for a more recent discussion. The pure Lifshitz geometry without a black brane develops singularities due to diverging tidal forces [18]; as a result the zero-temperature limit might become ill-defined¹. While this could be of possible concern for our analysis, our findings show that the fermionic Green's function is well defined at zero temperature and the limit $T \rightarrow 0$ is smooth, so at least the spectral-weight function for fermions does not suffer from any singularities.

In [14] we have described how to construct the retarded Green's function of a strongly interacting, but particle-hole symmetric system of chiral fermions with an arbitrary dynamical exponent z using a Lifshitz black-brane background. The aim of the present paper is to analyze the physics that follows from our prescription. In particular, we calculate the retarded single-particle Green's function in various cases, at zero and non-zero temperatures, and present the corresponding spectral-weight function and dispersion relations. Although we give some results for $z = 1$, we mostly focus on the case $z = 2$ and four spacetime dimensions. With this number of dimensions, the chiral fermions are Weyl fermions and the boundary system turns out to be an interacting Weyl semimetal² with Lifshitz scaling. By computing the momentum distribution of the particles and holes, we find for $z \neq 1$ that our holographic Weyl semimetal contains a quantum phase transition between a non-Fermi-liquid phase and a Fermi-liquid phase with two Fermi surfaces, one for the particles and one for the holes, even at zero doping. The scale set by the Fermi surface is determined by an appropriate combination of the dimensionless parameters λ and g that enter in our prescription for the Green's function on the

¹See however [19, 20] for possible resolutions within string theory.

²A semimetal is a gapless semiconductor. In addition, a Weyl semimetal [38] is a semimetal with touching valence and conduction bands based on chiral two-component fermions that in the non-interacting limit satisfy the Weyl equation $\pm \vec{\sigma} \cdot \vec{k} \psi = E \psi$. Here, the \pm denotes the chirality of the fermion.

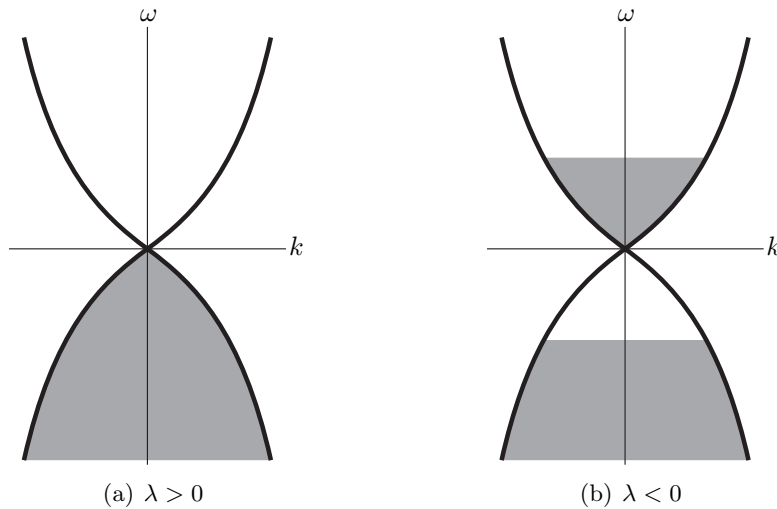


Figure 1: An idealized picture of the ground state that we find for $\lambda > 0$ and $\lambda < 0$. Grey areas refer to filled energy levels. In this picture, we only show the effect of the interactions on the occupation numbers and not on the dispersion relations.

holographic boundary [14], and the quantum critical point is at $\lambda = 0$. One can think of $1/\lambda$ as a spin-orbit coupling constant and of g as an effective interaction parameter coupling the elementary fermion to the conformal field theory. The phase transition is schematically illustrated in figure 1, where we show the non-interacting band structure and how it is populated in the ground state of each phase. In particular, for $\lambda > 0$ the conduction band is empty and the valence band is completely filled, whereas for $\lambda < 0$ the conduction band contains a Fermi sea of particles and the valence band a Fermi sea of holes. Note that we present here an idealized picture that neglects the renormalization of the Fermi surfaces involved, which is discussed at length in this paper.

The layout of the paper is as follows. In section 2, we briefly summarize some basic formulas from [14] that are needed here. In section 3, we present analytic and numerical results for the single-particle spectral function for the cases $z = 1$ and $z = 2$. In section 4 we consider the single-particle momentum distribution, and describe properties of the quantum phase transition and the Fermi liquid involved, that does allow for well-defined but strongly renormalized quasi-particles and quasi-holes near the two Fermi surfaces. The main part of the paper ends with conclusions in section 5. Finally, there are a number of appendices which contain our conventions and give more details about the calculations.

2 Single-particle Green's function for Lifshitz fermions

In this section, we determine solutions to the Dirac equation in a Lifshitz spacetime, from which we construct the retarded fermionic single-particle Green's function. To keep the paper self-contained, we start by defining our notation and conventions. Readers who

are only interested in the result of the calculation may skip directly to equation (2.11). Further details on the derivation can be found in the appendices, in particular, for remarks on how to switch from natural to SI units we would like to refer to appendix A.1. Some of the results in this section have already appeared in more detail in our previous work [14], other related work on fermions in Lifshitz backgrounds can be found in [21, 22, 23].

We begin by specifying the gravitational background metric of a Lifshitz black brane [16], i.e.,

$$ds^2 = -V^2(r) r^{2z} dt^2 + \frac{dr^2}{r^2 V^2(r)} + r^2 d\vec{x}^2, \quad V^2(r) = 1 - \left(\frac{r_h}{r}\right)^{d+z-1}, \quad (2.1)$$

where d denotes the number of spatial dimensions. The extra spatial coordinate r runs from the horizon at $r = r_h$ to the boundary at $r = \infty$, and the temperature of the black brane is obtained by demanding the absence of a conical singularity at r_h , leading to [16]

$$T = \frac{d+z-1}{4\pi} (r_h)^z. \quad (2.2)$$

The metric (2.1) enjoys the Lifshitz isometry, which will be inherited by the boundary theory as we will show shortly, and which reads

$$r \rightarrow \lambda r, \quad t \rightarrow \lambda^{-z} t, \quad x \rightarrow \lambda^{-1} x, \quad T \rightarrow \lambda^z T. \quad (2.3)$$

Next, we introduce a Dirac fermion Ψ into the above gravitational background. Following [11, 14], we include a boundary action on an ultra-violet (UV) cut-off surface at $r = r_0$. While many of the results can be phrased for arbitrary dimensions, for later purposes we specify to $d = 4$ spatial dimensions already here. In the conventions of [14], the total action for the bulk fermion that we consider is

$$S_{\text{total}}[\Psi] = i g_f \int d^5 x \sqrt{-g} \left(\frac{1}{2} \bar{\Psi} \vec{\mathcal{D}} \Psi - \frac{1}{2} \bar{\Psi} \overleftarrow{\mathcal{D}} \Psi - M \bar{\Psi} \Psi \right) - \int_{r=r_0} d^4 x \sqrt{-h} \left(\Psi_+^\dagger Z \mathcal{D}_z(r, x) \Psi_+ + i g_f \sqrt{g^{rr}} \Psi_+^\dagger \Psi_- \right), \quad (2.4)$$

where the first term on the right-hand side is the Dirac action in the bulk and the second term is the boundary action. Our notation is such that $\bar{\Psi} = \Psi^\dagger \Gamma^0$ with Γ^0 being anti-hermitian, and a natural choice for the Dirac matrices employed in this paper is given in appendix A.2. Furthermore, we have decomposed Ψ into chiral components according to its eigenvalue under Γ^z [8, 9] in the following way

$$\Psi \equiv \begin{pmatrix} \Psi_+ \\ \Psi_- \end{pmatrix}, \quad \Gamma^z \Psi = \begin{pmatrix} +\Psi_+ \\ -\Psi_- \end{pmatrix}. \quad (2.5)$$

The normalization constants g_f and Z appearing in (2.4) are left unspecified for the moment. The coupling of the fermions to the gravitational field is through the vielbeins $e_{\underline{a}}{}^\mu$ and the spin connection $(\Omega_\mu)_{\underline{a}\underline{b}}$, which are determined from the geometry and whose explicit form can be found in appendix A.2. The symbol \mathcal{D}_z appearing in the boundary action is the usual Dirac operator for a fermion in a Lifshitz background with arbitrary dynamic exponent z , and h denotes the determinant of the induced metric on the boundary.

Following [24], we also define Fourier-transformed spinors on each constant r slice as

$$\Psi_\pm(r, x) = \int \frac{d^4 p}{(2\pi)^4} \psi_\pm(r, p) e^{ip_\mu x^\mu}, \quad p_\mu = (-\omega, \vec{k}), \quad (2.6)$$

where ω and \vec{k} denote the frequency and momenta of the plane wave. The Dirac equation $(\mathcal{D} - M)\Psi = 0$ resulting from (2.4) (see [14] for more details) and in-falling boundary conditions that we shall employ at the horizon imply a relation between ψ_+ and ψ_- . It can be expressed as

$$\psi_-(r, p) = -i\xi(r, p) \psi_+(r, p), \quad (2.7)$$

and can be used to integrate out ψ_- from the action. Together with (2.4) and (2.6) this results in a holographic effective action for the field ψ_+ on the cut-off surface as

$$S_{\text{eff}}[\Psi_+] = - \int_{r=r_0} \frac{d^4 p}{(2\pi)^4} \sqrt{-h} \psi_+^\dagger [Z\mathcal{D}_z(p) + g_f \sqrt{g^{rr}} \xi(r_0, p)] \psi_+. \quad (2.8)$$

The Green's function of ψ_+ that follows from (2.8) for our geometry (2.1) reads

$$G_R(r_0, p) = - \left(r_0^z V(r_0) \mathcal{D}_z(p) + \frac{g_f}{Z} r_0^{1+z} V^2(r_0) \xi(r_0, p) \right)^{-1}, \quad (2.9)$$

where we have rescaled ψ_+ so that it acquires a canonically normalized kinetic term. The final step is now to take a double-scaling limit

$$r_0 \rightarrow \infty, \quad g_f \rightarrow 0, \quad g_f r_0^{1+z-2M} = \text{const.} \quad (2.10)$$

As we take this limit, we make sure that the resulting effective action has a kinetic term expected from a theory with dynamical scaling z . In the case $z \neq 1$, this can be achieved by renormalizing away a relativistic kinetic term and adding appropriate counter-terms to the action [14]. The result for $z \neq 1$ then is

$$G_R(\vec{k}, \omega) = - \left(\omega - \frac{1}{\lambda} \vec{\sigma} \cdot \vec{k} |\vec{k}|^{z-1} - \Sigma(p) \right)^{-1}, \quad (2.11)$$

where λ is an arbitrary real number in our approach,³ and where in the following we denote $k = |\vec{k}|$. The above formula for the retarded Green's function is also valid for

³In the notation of [14], the parameter λ appears as $\lambda = -1/\eta$.

$z = 1$, but then the relativistic invariance requires that $\lambda = 1$. The self-energy $\Sigma(p)$ is the interaction term that arises from coupling the elementary field ψ_+ to the conformal field theory encoded in the Lifshitz background. As can be seen by comparing with (2.9), it is related to the quantity $\xi(r, p)$ introduced in (2.7) by

$$\Sigma(p) = -g \lim_{r_0 \rightarrow \infty} r_0^{2M} \xi(r_0, p), \quad -\frac{1}{2} < M < \frac{1}{2}, \quad (2.12)$$

where g is the coupling constant which stays finite in the limit (2.10) and reads

$$g = \frac{gf}{Z} r_0^{1+z-2M}. \quad (2.13)$$

The definition of the self-energy $\Sigma(p)$ in (2.12) is valid for all values of the momentum \vec{k} , however, for $\vec{k} = 0$ the allowed range of M is extended to $-z/2 < M < z/2$. This can be derived from the asymptotic behavior of $\xi(r, p)$ near the boundary, and for more details we refer the reader to appendix B. Furthermore, it is possible to extend the relation between $\xi(r, p)$ and $\Sigma(p)$ even to $|M| > 1/2$ for non-zero momentum, which requires introducing certain counter-terms on the cut-off surface at r_0 before taking the limit $r_0 \rightarrow \infty$. Since we will also be interested in the range $1/2 < M < z/2$, using the results in appendix B, we can show that the relation (2.12) in this case should be modified to

$$\Sigma(p) = -g \lim_{r_0 \rightarrow \infty} \left(r_0^{2M} \xi(r_0, p) - \frac{\vec{\sigma} \cdot \vec{k}}{2M - 1} r_0^{2M-1} \right), \quad \frac{1}{2} < M < \frac{z}{2}. \quad (2.14)$$

Note that the second term in this expression, which is divergent for $M > 1/2$, removes the divergence in $\xi(r, p)$ and yields a finite result for $\Sigma(p)$.

The transfer matrix $\xi(r, p)$ defined in equation (2.7) is a complex two-by-two matrix. In the case we are interested in, that is, $d = 4$, it can be diagonalized by choosing the Weyl basis of gamma matrices (A.15). A first-order differential equation for the eigenvalues $\xi_{\pm}(r, p)$ of $\xi(r, p)$ can be derived, which was achieved in the anti-de-Sitter case ($z = 1$) in [24] and generalized to arbitrary z in [14]. The resulting differential equations read

$$r^2 V \partial_r \xi_{\pm} + 2Mr \xi_{\pm} = \frac{\omega}{r^{z-1}V} \mp k_3 + \left(\frac{\omega}{r^{z-1}V} \pm k_3 \right) \xi_{\pm}^2, \quad (2.15)$$

where we used rotational invariance to set $\vec{k} = (0, 0, k_3)$. Imposing in-falling boundary conditions at the black-brane horizon, which corresponds to considering the retarded Green's function, leads to the boundary condition

$$\xi_{\pm}(r_h) = i. \quad (2.16)$$

The functions $\xi_{\pm}(r, p)$ as well as the resulting self-energy $\Sigma(p)$ have various symmetries, which are discussed in more detail in appendix C.

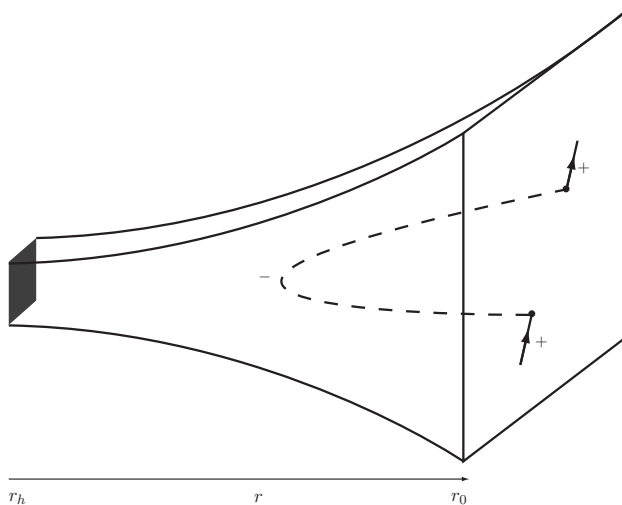


Figure 2: Illustration of our holographic construction of the self-energy.

From the imaginary part of the retarded Green's function G_R , the total spectral weight, or spectral function, can be obtained, which is of great importance in condensed-matter physics and is also directly observable in experiments. It is defined as the trace of the two-by-two matrix G_R , i.e.,

$$\rho(\vec{k}, \omega) = \frac{1}{2\pi} \text{Im Tr} [G_R(\vec{k}, \omega)] . \quad (2.17)$$

It was shown in [14] that the spectral-weight function corresponding to (2.11) satisfies the sum rule

$$\int_{-\infty}^{\infty} d\omega \rho(\vec{k}, \omega) = 1 \quad (2.18)$$

in the physically allowed range for the scale dimensions of the CFT operator coupled to ψ . In [14], this range was found to be $-z/2 < M < z/2$, where the lower bound is due to unitarity in the field theory, and the upper bound is a consequence of the requirement that the coupling of the elementary fermion to the CFT is irrelevant in the ultra-violet. Within this allowed physical range for M , the Kramers-Kronig relations are satisfied and the sum rule (2.18) is obeyed. The latter is a consequence of the fact that G_R is the Green's function for an elementary field ψ and its hermitian conjugate ψ^\dagger , that satisfy anti-commutation relations. In conclusion, the prescription (2.11) thus allows us to holographically compute *single-particle* Green's functions, which are important from a condensed-matter perspective.

To close this section, we summarize and illustrate our holographic construction in figure 2. In particular, equation (2.11) describes chiral single-particle propagators on the boundary of the spacetime. These chiral fermions interact with a conformal field theory due to a fermion of opposite chirality that travels into the bulk spacetime, has classical

interactions there, and comes back to the boundary. In the ultra-violet, the points where the chiral fermion emits and reabsorbs the fermion of opposite chirality are very close together and the latter fermion cannot travel far into the bulk. Hence, it only notices the flat background on the cut-off surface in the bulk spacetime, which leads to a free fermionic propagator in the UV. For the infra-red (IR) case, however, the fermion of opposite chirality can indeed travel far into the curved part of the bulk spacetime, and as a result the IR dynamics will be dominated by interactions and will generally not be free. This corresponds physically to the usual “energy-scale” interpretation of the extra spatial direction.

3 Single-particle spectra

In the previous section, we have outlined the construction of the retarded single-particle Green’s function. In the present section, we now study this Green’s function for both the relativistic $z = 1$ and non-relativistic $z = 2$ case in more detail, and work out its physical properties. The zero-temperature results were first obtained in [14], and here we consider non-zero temperatures as well. However, let us stress that the systems we describe are always at zero chemical potential.

3.1 Relativistic case $z = 1$

In the case that the elementary fermion ψ_+ interacts with a relativistic CFT, the background is given by an AdS black brane and is described by the metric shown in equation (2.1) for $z = 1$. The mass range for the bulk Dirac fermion is then restricted to lie within the interval $-1/2 < M < 1/2$.

3.1.1 Zero temperature

For vanishing temperature and $z = 1$, the self-energy $\Sigma(p)$ can be computed analytically for arbitrary frequencies ω and momenta \vec{k} . The result for the full retarded Green’s function has been worked out in [14] and reads

$$G_R(\vec{k}, \omega) = -\frac{1}{p^2 \left(1 - g c_1 e^{-i\pi(M+\frac{1}{2})} p^{2M-1}\right)} \left(\omega + \vec{\sigma} \cdot \vec{k}\right), \quad (3.1)$$

where we defined $p \equiv \sqrt{\omega^2 - \vec{k} \cdot \vec{k}}$ as well as the constant (for arbitrary values of z)

$$c_z = (2z)^{-\frac{2M}{z}} \frac{\Gamma\left(\frac{1}{2} - \frac{M}{z}\right)}{\Gamma\left(\frac{1}{2} + \frac{M}{z}\right)}. \quad (3.2)$$

Note that (3.1) can be extended into the complex plane by allowing for complex momenta p , which is important when determining the pole structure of the Green’s function. Because of the branch point at $p = 0$, we need to introduce a branch cut which is taken to run from $p = 0$ to $p = -i\infty$ for later convenience. Furthermore, we find that in

order for (3.1) to be free of singularities in the upper half ω plane, i.e., to satisfy the Kramers-Kronig relations, we have to demand

$$g > 0, \quad (3.3)$$

which is derived in detail in appendix D.1. It is then straightforward to show that the Green's function (3.1) also satisfies the sum rule (2.18), as it was done in [14] and summarized in appendices D.2 and D.3. Furthermore, as mentioned above, (3.1) is valid on the complex p plane, with the prescription that on the real ω line the self-energy is found by using

$$p^{2M-1}e^{-i\pi(M+\frac{1}{2})} = \begin{cases} +p^{2M-1}e^{-i\pi(M+\frac{1}{2})}, & p \equiv \sqrt{\omega^2 - \vec{k} \cdot \vec{k}}, & \omega > +|\vec{k}|, \\ +p^{2M-1}e^{+i\pi(M+\frac{1}{2})}, & p \equiv \sqrt{\omega^2 - \vec{k} \cdot \vec{k}}, & \omega < -|\vec{k}|, \\ -p^{2M-1}, & p \equiv \sqrt{\vec{k} \cdot \vec{k} - \omega^2}, & -|\vec{k}| < \omega < +|\vec{k}|. \end{cases} \quad (3.4)$$

This directly follows from (3.1) by noting that the region $\omega < -|\vec{k}|$ for real frequencies is obtained from the region $\omega > |\vec{k}|$ by $p \rightarrow e^{i\pi}p$, that is $\omega \rightarrow e^{i\pi}\omega$ and $\vec{k} \rightarrow e^{i\pi}\vec{k}$, while the region $|\omega| < |\vec{k}|$ is obtained via $p \rightarrow e^{i\pi/2}p$. Note also that the Green's function for $-|\vec{k}| < \omega < +|\vec{k}|$ has no imaginary part.

3.1.2 Non-zero temperature

We have studied the effects of non-zero temperature on the structure and form of the Green's function by numerical integration of the Dirac equation. To illustrate our results, we have included figures 3 and 4 which show the total spectral-weight function (2.17) for $T = 1/30$ and $T = 2$, respectively. Furthermore, we made a distinction between positive and negative masses M , and we employed the rotational symmetry to set $\vec{k} = (0, 0, k_3)$. Let us discuss these figures in some more detail:

- First, we note that a non-vanishing temperature results in a smearing out of the features of the Green's function, as can be seen by comparing figures 3 and 4.
- For $M > 0$ we observe a large spectral weight approximately at $\omega = \pm k_3$, which is due to the “light cone” of relativistic physics. In the zero-temperature case (3.4), the imaginary part of the Green's function is strictly zero outside of the light cone in the region $-|\vec{k}| < \omega < +|\vec{k}|$. For non-zero temperatures, there is a small contribution in this region, which increases with temperature.

Note that the peak in the spectral weight is not caused by a pole in the Green's function, but by the aforementioned branch cut. Therefore, it has no interpretation as a well-defined quasi-particle excitation, but it is still possible to approximately determine the dispersion relation, as it is shown in the next subsection.

- For negative masses $M < 0$, the spectrum has a similar form. Indeed, there is again a large peak at $\omega = \pm k_3$. However, for low T , the imaginary part of the

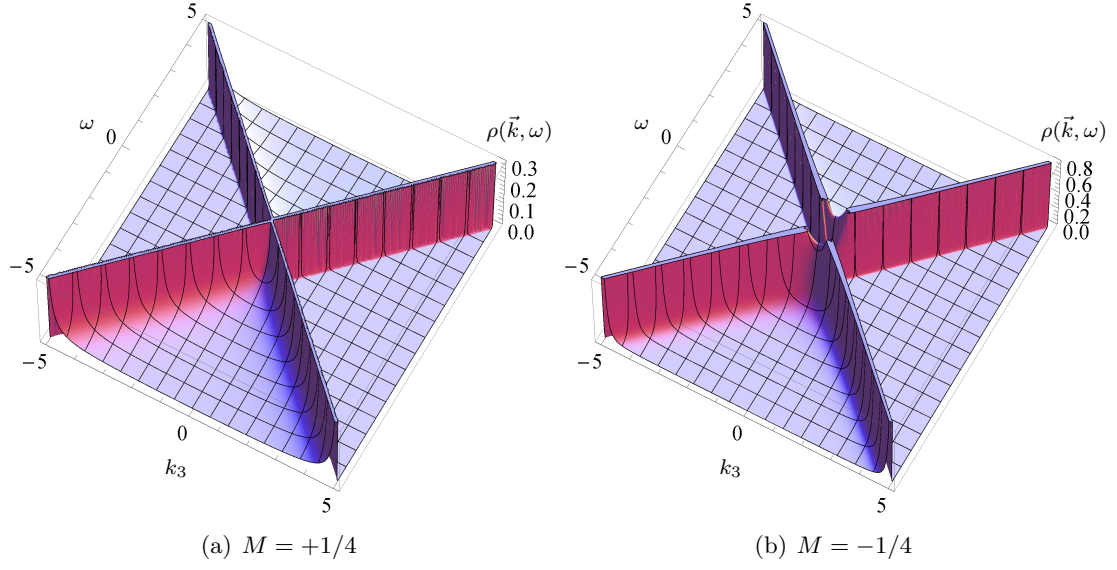


Figure 3: Total spectral function for $T = 1/30$, $g = 1$ and $M = \pm 1/4$.

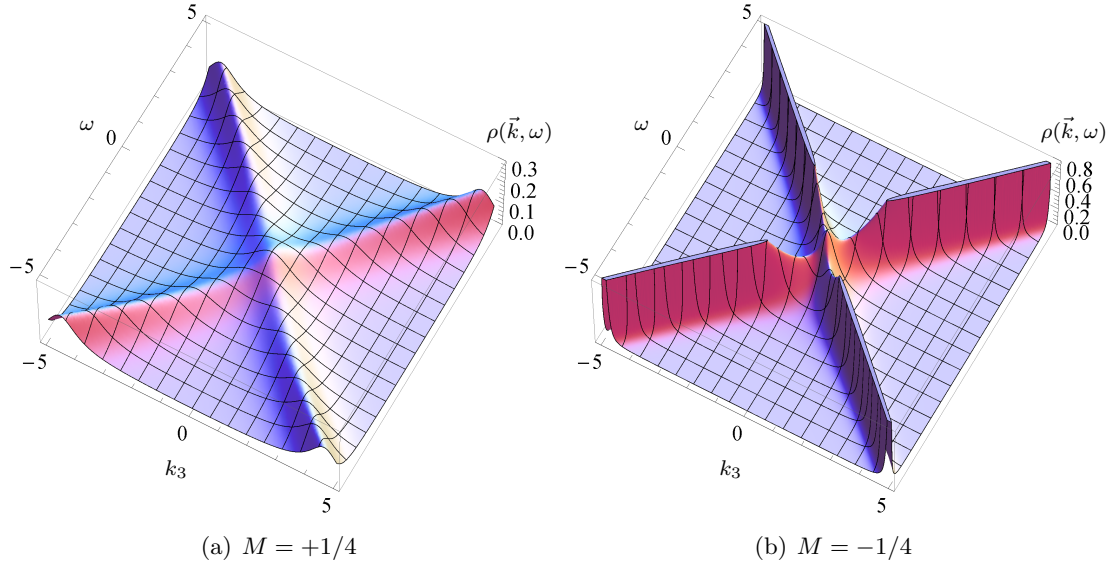


Figure 4: Total spectral function for $T = 2$, $g = 1$ and $M = \pm 1/4$.

Green's function goes to zero at $k_3 = 0$ as $\omega \rightarrow 0$. It has a maximum in between, which leads to a broad maximum in the spectrum. Its approximate location is determined analytically in section 3.1.3.

We have also investigated the ratio of the location of the maximum and its width as a function of g . This ratio is approximately constant, which means that its shape becomes broader as we increase its position, and vice versa. The spectral function at $k_3 = 0$ looks very similar to what is plotted in figure 7 for the non-relativistic case with $z = 2$. It is not possible to tune the parameters such that the peak is located sufficiently far away from the origin at $\omega = 0$ while its maximum remains sharp. Therefore, we cannot interpret this feature in the spectral function as a massive particle. Again, we give some further analytic arguments of this statement at zero temperature in section 3.1.3.

3.1.3 Dispersion relation

In the last subsection 3.1.2, we have illustrated that the spectral weight is peaked at particular functions $\omega(\vec{k})$. In this section, we now investigate the corresponding dispersion relation of the theory. For vanishing temperature this can be done analytically via the single-particle Green's function by solving

$$\text{Re } G_R^{-1}(\vec{k}, \omega) = 0 . \quad (3.5)$$

By definition, the solutions to this equation determine the dispersion relation of the would-be (quasi-)particle. The interpretation as a particle only becomes justified if the width is small compared to the energy of the particle. Indeed, a large peak in the spectral density function is only obtained when both $\text{Re}(G_R^{-1})$ and $(\text{Im } \Sigma)^2$ are minimal, as follows from the identity

$$\text{Im } G_R = - \frac{\text{Im } \Sigma}{(\text{Re } G_R^{-1})^2 + (\text{Im } \Sigma)^2} . \quad (3.6)$$

Let us first solve (3.5). Using the explicit expression (3.1), equation (3.5) generically gives two possible dispersion relations, namely

$$\omega = \pm |\vec{k}| \quad \text{and} \quad g c_1 \text{Re} \left[e^{-i\pi(M+\frac{1}{2})} p^{2M-1} \right] = 1 . \quad (3.7)$$

The first solution describes a free, massless relativistic excitation which is always present. The second equation in (3.7) should be studied separately in the two regions $|\omega| < |\vec{k}|$ and $|\omega| > |\vec{k}|$. In the first case, using (3.4) we obtain the solution

$$p^2 = \left(-\frac{1}{g c_1} \right)^{\frac{2}{2M-1}} . \quad (3.8)$$

Since $g c_1 > 0$ and the left-hand side of (3.8) should be real, we conclude that there is no solution for $|\omega| < |\vec{k}|$. In the second case, $|\omega| > |\vec{k}|$, the solution reads

$$p^2 = \left(g c_1 \cos \left[\pi \left(M + \frac{1}{2} \right) \right] \right)^{\frac{2}{1-2M}} . \quad (3.9)$$

For $0 < M < 1/2$, we do not find a solution since $g c_1 > 0$. On the other hand, for $-1/2 < M < 0$, we obtain

$$\omega = \pm \sqrt{\vec{k} \cdot \vec{k} + m^2}, \quad m^2 = \left(g c_1 \cos \left[\pi \left(M + \frac{1}{2} \right) \right] \right)^{\frac{2}{1-2M}}, \quad (3.10)$$

which is indeed close to the locations of the maxima in the spectral-weight function that were found for non-zero temperature. For zero temperature, equation (3.10) yields the exact result.

Now we can look at the width, by computing the imaginary part of the self-energy at the values (3.10). A straightforward calculation shows that, at zero temperature, we have

$$\text{Im } \Sigma = -(\omega - \vec{\sigma} \cdot \vec{k}) \tan \left[\pi \left(M + \frac{1}{2} \right) \right], \quad (3.11)$$

with ω given by (3.10). However, for momenta small compared to the gap m (in the restframe of the would-be particle), we can approximate

$$\text{Im } \Sigma \approx m \tan \left[\pi \left(M + \frac{1}{2} \right) \right]. \quad (3.12)$$

The width is then comparable to the gap, and therefore the peak in the spectral density does not have a quasi-particle interpretation. One might try to make the width smaller by taking the value of M close to the unitarity bound, $M = -1/2$. However, notice that then the gap also narrows down. We conclude therefore that no true quasi-particles exist.

3.2 Non-relativistic case $z \neq 1$

In principle, every value of the dynamical exponent z can be considered using the prescription shown in (2.11). However, for our purposes we are mostly interested in $z = 2$ where the elementary fermion interacts with a CFT exhibiting $z = 2$ Lifshitz scale invariance.

3.2.1 Zero temperature

Unlike the relativistic case studied above, for an arbitrary dynamical exponent z we were not able to obtain $G_R(\omega, \vec{k})$ analytically for the both ω and \vec{k} non-vanishing. But it is possible to determine the zero-temperature result for $\vec{k} = 0$ and $\omega = 0$ separately, and then restrict the general expression to a great extent.

Let us therefore start with the case $\vec{k} = 0$ and $T = 0$. From [14] we recall the expression for the Green's function as

$$G_R(\vec{0}, \omega) = -\frac{1}{\omega - g c_z \omega^{\frac{2M}{z}} e^{-i\pi(\frac{M}{z} + \frac{1}{2})}}, \quad (3.13)$$

where we employed the definition (3.2). In the case with $\omega = 0$ and $T = 0$ we instead find [14]

$$G_R(\vec{k}, 0) = \frac{1}{\frac{1}{\lambda} \vec{\sigma} \cdot \vec{k} k^{z-1} + g c_1 k^{2M-1} \vec{\sigma} \cdot \vec{k}}, \quad (3.14)$$

where again k denotes $|\vec{k}|$. The generic case at vanishing temperature when both \vec{k} and ω are non-zero can then be restricted as follows. Using the scaling and rotational symmetries of the self-energy (see appendix C.1), and defining $u = \omega/k^z$ for notational simplicity, we have

$$G_R(\vec{k}, \omega) = -\frac{1}{\omega - \frac{1}{\lambda} \vec{\sigma} \cdot \vec{k} k^{z-1} - g \left(k^{2M} s_{1,M}(u) + k^{2M-1} s_{2,M}(u) \vec{\sigma} \cdot \vec{k} \right)}, \quad (3.15)$$

where $s_{1,M}$ and $s_{2,M}$ are complex functions of u and M only. Furthermore, we can derive conditions on the functions $s_{1,2}$ using the symmetry of $\Sigma(p)$ under the change $M \rightarrow -M$, as discussed in appendix C.4. The asymptotic behavior of the functions $s_{1,M}(u)$ and $s_{2,M}(u)$ in the limits $u \rightarrow 0$ and $u \rightarrow \infty$ is fixed by the expressions (3.13) and (3.14) given above. We therefore have

$$\begin{aligned} s_{1,M}(u) &\xrightarrow{u \rightarrow \infty} u^{\frac{2M}{z}} e^{-i\pi(\frac{M}{z} + \frac{1}{2})} c_z, & s_{2,M}(u) &\xrightarrow{u \rightarrow \infty} 0, \\ s_{1,M}(u) &\xrightarrow{u \rightarrow 0} 0, & s_{2,M}(u) &\xrightarrow{u \rightarrow 0} c_1. \end{aligned} \quad (3.16)$$

Unfortunately, these conditions do not seem to be sufficient to determine the analytic form of (3.15). Therefore we have to study the Green's function numerically. Yet, we can determine the qualitative form of the dispersion relation by analytic arguments as we describe below in section 3.2.3.

3.2.2 Non-zero temperature

We have studied the retarded Green's function and the corresponding spectral-weight function for $z = 2$ numerically as a function of ω , \vec{k} , T and λ . Using the symmetries summarized in appendix C, that is chirality and particle-hole symmetry, we observe that the components of the spectral-weight function obey the following relations

$$\rho^\pm(\vec{k}, \omega) = \rho^\mp(-\vec{k}, \omega) = \rho^\pm(-\vec{k}, -\omega). \quad (3.17)$$

Consequently, it suffices to consider separate components instead of the trace over chiral components.

Next, we recall that for $z \neq 1$ the Green's function (2.11) contains the parameter λ , which we have not determined analytically. Treating λ as a free parameter, we observe that under a change of sign of λ , the far UV behavior of one component of the spectral weight function asymptotes the UV behavior of the other component with the original sign, that is,

$$\rho_{-\lambda}^\pm(\vec{k}, \omega) \xrightarrow{|\omega| \rightarrow \infty, |\vec{k}| \rightarrow \infty} \rho_{+\lambda}^\mp(\vec{k}, \omega) = \rho_{+\lambda}^\pm(-\vec{k}, \omega). \quad (3.18)$$

Since λ can be interpreted as a spin-orbit coupling constant, we use the following convention for plotting components of the spectral-weight function: when λ is positive (negative), the plus(minus) component of the spectral-weight function is shown. In this manner we make sure that we always compare spectra with equal group velocity in the UV, also when λ changes sign. We make this choice because qualitatively the UV physics is then, apart from the topology of the band structure, independent of the sign of λ . This allows us to compare more clearly the physics for positive and negative values of λ as we will see shortly.

In figures 5 and 6 we show numerical results for separate components of the corresponding spectral-weight function. The parameters are chosen as $M = \pm 1/4$, $g = 1$ and $T = 1/30$, for a number of different values for λ . We discuss some of the features in turn:

- In figure 5, taken for positive M , we see that the spectral-weight function is very sharply peaked in the UV and behaves as a free chiral fermion with a quadratic dispersion. In the IR, the self-energy becomes dominant which changes and smears out the form of the free-fermion dispersion. When λ switches sign, the spectral-weight function changes in the IR. However, note that our convention is such that the UV behavior stays the same.
- In figure 6, where M is negative, we observe a phenomenon similar to the relativistic case with $M < 0$. In particular, the spectral weight has to vanish for $k = 0$ as $\omega \rightarrow 0$, so there is a maximum in the spectral-weight function at non-zero ω . However, the ratio of the location and the width of the peak again remains constant as we change g , so there is no gap generation with a quasi-particle interpretation. This is further illustrated in figure 7.

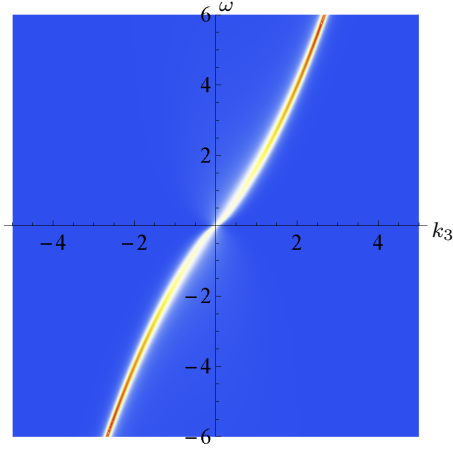
To close our discussion about non-zero temperature effects in the case of $z = 2$, let us consider the strict IR (or hydrodynamic) limit $\omega \rightarrow 0$, $k \rightarrow 0$. As explained in appendix B.3, in this limit the contribution from the free propagator vanishes and the Green's function reduces to the inverse of the self-energy $\Sigma(p)$. An analytical result for all values of the dynamical exponent z and dimension d and for non-zero temperature has been obtained for this case in equation (B.21), which we recall here for convenience

$$G_R(\vec{0}, 0) = \frac{i}{g} 2^{\frac{4M}{d+z-1}} \left(\frac{d+z-1}{4\pi} \right)^{\frac{2M}{z}} T^{-\frac{2M}{z}}. \quad (3.19)$$

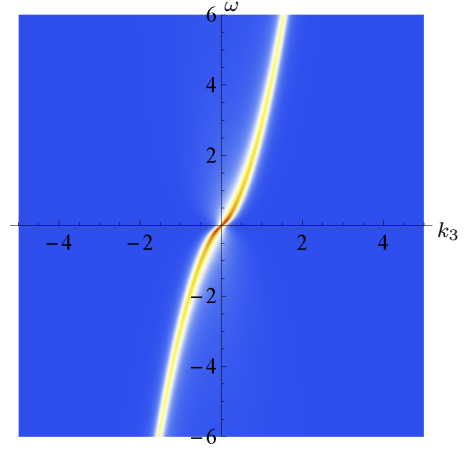
3.2.3 Dispersion relation

We now consider the dispersion relation for $z \neq 1$ at zero temperature, which can be derived from the general form of the Green's function given in (3.15). In particular, the dispersion relation is obtained by solving (3.5), which in the present case reads

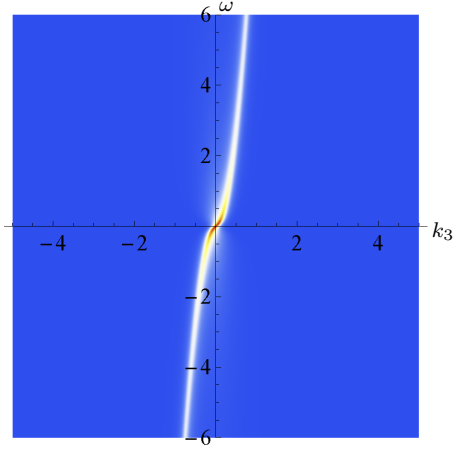
$$\omega - \frac{1}{\lambda} \vec{\sigma} \cdot \vec{k} k^{z-1} - g k^{2M} \text{Re} \left(s_{1,M}(u) + \frac{\vec{\sigma} \cdot \vec{k}}{k} s_{2,M}(u) \right) = 0. \quad (3.20)$$



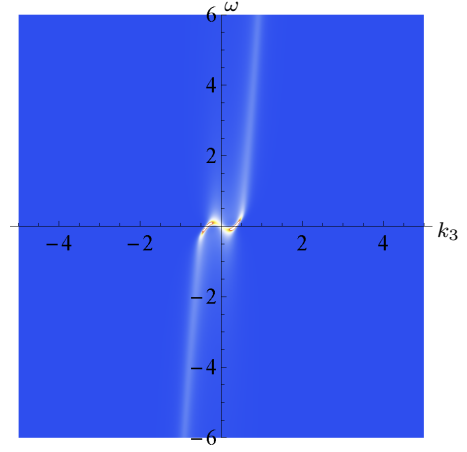
(a) Density plot of $\rho^+(\vec{k}, \omega)$ at $\lambda = 2$



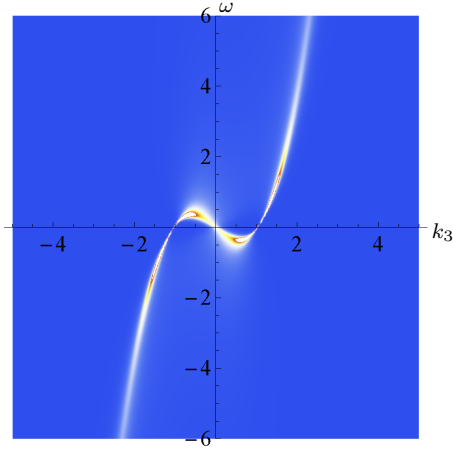
(b) Density plot of $\rho^+(\vec{k}, \omega)$ at $\lambda = 0.5$



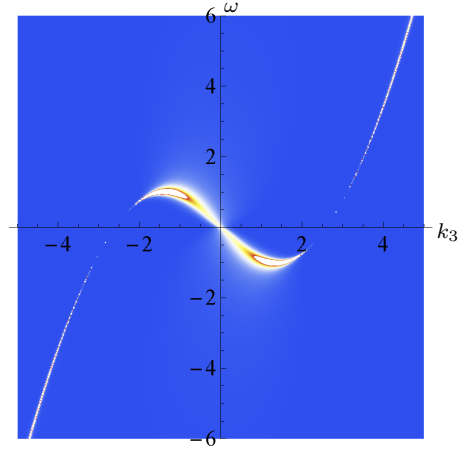
(c) Density plot of $\rho^+(\vec{k}, \omega)$ at $\lambda = 0.1$



(d) Density plot of $\rho^-(\vec{k}, \omega)$ at $\lambda = -0.1$

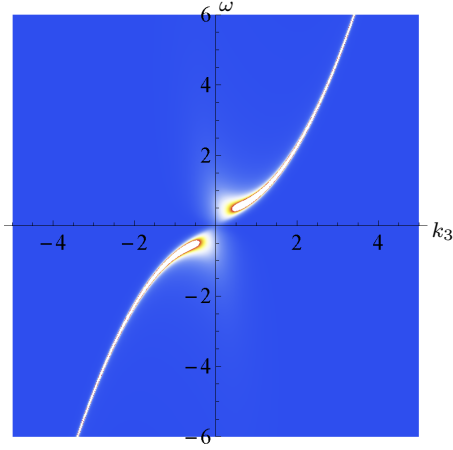


(e) Density plot of $\rho^-(\vec{k}, \omega)$ at $\lambda = -0.5$

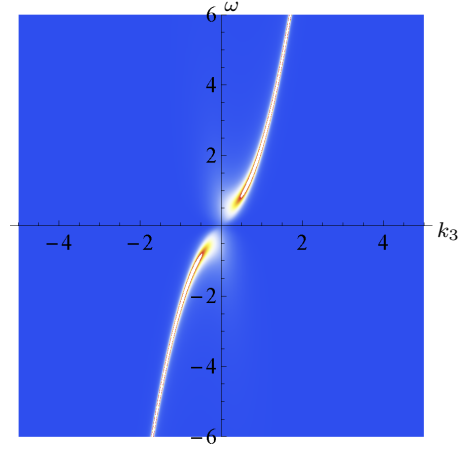


(f) Density plot of $\rho^-(\vec{k}, \omega)$ at $\lambda = -2$

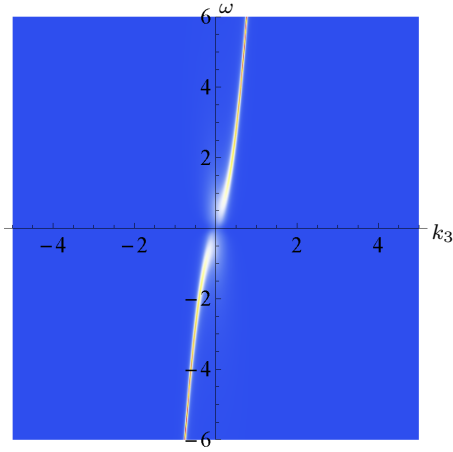
Figure 5: Density plot of $\rho^\pm(\vec{k}, \omega)$ for $z = 2$, $M = +1/4$, $T = 1/30$, $g = 1$.



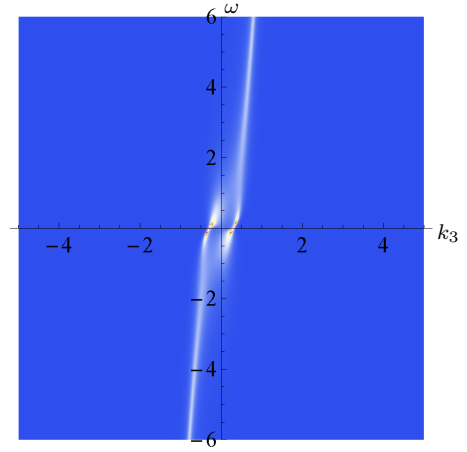
(a) Density plot of $\rho^+(\vec{k}, \omega)$ at $\lambda = 2$



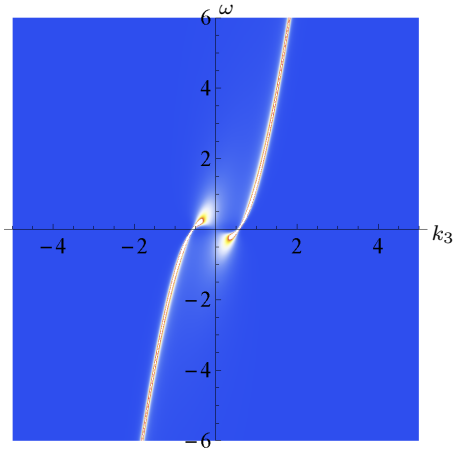
(b) Density plot of $\rho^+(\vec{k}, \omega)$ at $\lambda = 0.5$



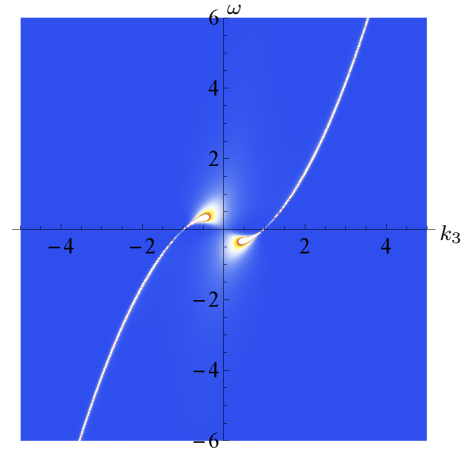
(c) Density plot of $\rho^+(\vec{k}, \omega)$ at $\lambda = 0.1$



(d) Density plot of $\rho^-(\vec{k}, \omega)$ at $\lambda = -0.1$



(e) Density plot of $\rho^-(\vec{k}, \omega)$ at $\lambda = -0.5$



(f) Density plot of $\rho^-(\vec{k}, \omega)$ at $\lambda = -2$

Figure 6: Density plot of $\rho^\pm(\vec{k}, \omega)$ for $z = 2$, $M = -1/4$, $T = 1/30$, $g = 1$.

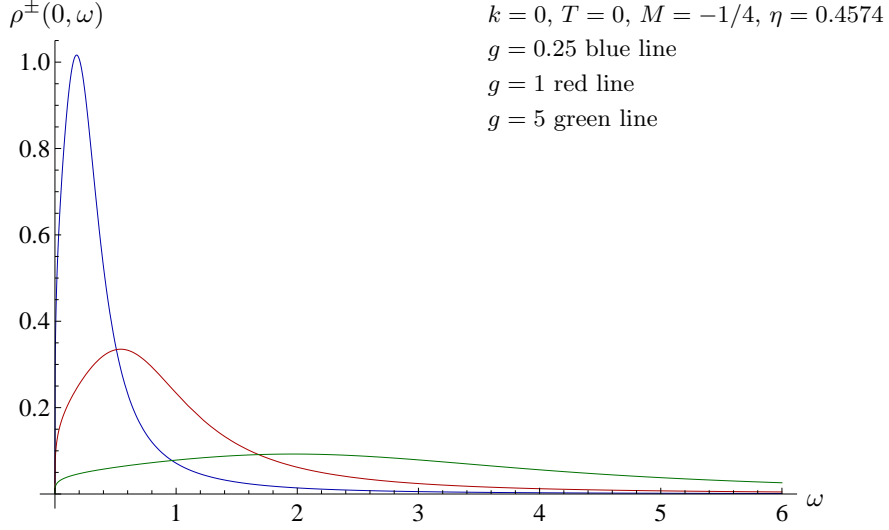


Figure 7: Spectral-weight function for vanishing momentum determined from (3.13). Since the ratio of the location and the width of the peak remains constant, there is no quasi-particle interpretation.

In the following we consider the upper-spin component for definiteness. The latter means that $\vec{\sigma} \cdot \vec{k}/k$ is replaced by $\text{sign}(k_3)$, where we again employed the rotational symmetry to align the momenta in the z -direction. For positive (negative) momentum k_3 the dispersion relation then becomes

$$\omega \mp \frac{1}{\lambda} k^z - g k^{2M} \text{Re}(s_{1,M}(u) \pm s_{2,M}(u)) = 0. \quad (3.21)$$

In order to determine the qualitative form of the dispersion relation, we study various limits of the equation above.

- We first consider the UV limit $\omega \rightarrow \infty$, $k \rightarrow \infty$ for which there are the three distinct possibilities: $\omega/k^z \rightarrow \infty$, $k^z/\omega \rightarrow \infty$ and $\omega/k^z \rightarrow \text{const.}$ The first two do not allow for a solution of (3.21), but the third possibility leads to

$$\omega \approx \pm \frac{1}{\lambda} k^z \quad \text{for } k \gg 1 \quad (3.22)$$

in the allowed mass range $-z/2 < M < z/2$. This result confirms our general picture that the interaction of the elementary fermion with the CFT is irrelevant in the UV, hence the dispersion becomes that of a free fermion.

- Next, we consider the limit $k_3 \rightarrow 0^+$, $\omega > 0$. In the mass range $-z/2 < M < z/2$ we determine the dispersion curve from (3.13) as

$$\omega + g c_z \left| \cos \left[\pi \left(\frac{M}{z} + \frac{1}{2} \right) \right] \right| \text{sign}(M) \omega^{\frac{2M}{z}} = 0. \quad (3.23)$$

Noting that $g > 0$ and $c_z > 0$, we find that the only solution to this equation is $\omega = 0^+$ for $M > 0$. On the other hand, for $M < 0$ we obtain the only solution as

$$\omega = \left(g c_z \left| \cos \left[\pi \left(\frac{M}{z} + \frac{1}{2} \right) \right] \right| \right)^{\frac{z}{z-2M}} \equiv m_z, \quad M < 0. \quad (3.24)$$

These solutions correspond to the points where the dispersion curve touches the axis at $k = 0$.

Studying the dispersion relation for the case of $k_3 \rightarrow 0^+$ and $\omega < 0$, we find another solution at $\omega = -m_z$, consistent with the particle-hole symmetry. This is the non-relativistic analog of the result we obtained in equation (3.10). As mentioned above, and similar to the case with $z = 1$, the corresponding peaks in the spectral-weight function cannot be interpreted as a true gapped quasi-particle excitation. This is further illustrated by a plot of the analytic result in figure 7.

- The most interesting case is when $0 < M < 1/2$ and $\lambda < 0$. As one can see from figure 5, this situation is quite different from $0 < M < 1/2$ and $\lambda > 0$. The dispersion curve $\omega(\vec{k})$ for $\lambda < 0$ crosses the $\omega = 0$ axis three times, instead of only once at $\omega = k = 0$. It therefore seems that two Fermi surfaces appear at the momenta for which $\omega(\pm k_F) = 0$. This signals a phase transition as one crosses from positive to negative λ . The remainder of this paper is devoted to a detailed study of this phase transition.

4 Quantum phase transition

In figures 5 and 6 we have already seen that for $z = 2$ the parameter λ controls features of the spectral density significantly. In particular, when switching the sign of λ the number of zeros of the dispersion relation $\omega(\vec{k})$ changes, as becomes clear when comparing for instance figures 5(a) and 5(f). In the present section, we study this phenomenon in more detail. For simplicity we focus on the case $0 < M < 1/2$ since the quantum phase transition discussed here is also present for $-1/2 < M < 0$. However, we also briefly consider the case $|M| > 1/2$ in section 4.5.

4.1 Momentum distribution

Let us start by defining the momentum distribution function. We choose again a convention such that the group velocity of a particle with one of the spin components always has the same sign at large momenta, irrespectively of the sign of λ . More concretely, with $n_F(\omega)$ the Fermi distribution we write

$$N_k^\pm = \begin{cases} \frac{1}{\pi} \int_{-\infty}^{\infty} d\omega \operatorname{Im} G_R^\pm(\omega, \vec{k}) n_F(\omega) & \text{for } \lambda > 0, \\ \frac{1}{\pi} \int_{-\infty}^{\infty} d\omega \operatorname{Im} G_R^\mp(\omega, \vec{k}) n_F(\omega) & \text{for } \lambda < 0. \end{cases} \quad (4.1)$$

The behavior of $N_{\vec{k}}^{\pm}$ when λ changes sign is illustrated in figure 8 for non-zero temperatures, where we again employed the rotational symmetry to set $\vec{k} = (0, 0, k_3)$. In particular, for $\lambda < 0$ the momentum distribution indicates a Fermi surface with a certain width, on which we comment later. When λ switches sign, two extrema appear that develop into sharp discontinuities as λ increases. This is a clear signature of a Fermi surface. The locations of the jumps are determined analytically in the next section.

4.2 Fermi momentum

We now study analytically how the number of putative Fermi surfaces changes as we vary the parameter λ . To do so, we recall that Fermi surfaces are determined by the poles of the Green's function at vanishing frequencies $\omega = 0$. The latter can be computed by setting to zero the denominator in (2.11). Note also that the self-energy is given by (2.12) which satisfies the particle-hole symmetry derived from (C.16). Employing the general form of the self-energy implied by (3.15), we see that in the limit $\omega \rightarrow 0$ the imaginary part of $\Sigma(p)$ vanishes and the condition for the presence of a Fermi surface coincides with the presence of zeros of the dispersion relation (3.5). Thus, the loci of the Fermi surfaces are given precisely by the points where the dispersion curve $\omega(\vec{k})$ crosses the $\omega = 0$ axis.

We therefore consider equation (3.21) in the limit $\omega \rightarrow 0$. Using (3.16) and the expression (3.2) for c_1 , we obtain the following formula for the loci of the Fermi surfaces

$$\frac{1}{\lambda} k^z + g k^{2M} c_1 = 0. \quad (4.2)$$

Then, as we have mentioned in equation (3.3), to avoid violation of causality we have to require $g > 0$. Equation (4.2) can thus have non-trivial solutions in the range $-1/2 < M < 1/2$ only when $\lambda < 0$, that is

$$k = k_F = (-g\lambda c_1)^{\frac{1}{z-2M}}. \quad (4.3)$$

Next, let us consider again the numerical results for the imaginary part of the Green's function shown in figure 5. For $\lambda < 0$, there are zero-energy modes at non-vanishing momentum, suggesting that the system indeed has two Fermi surfaces at zero temperature. However, due to the small but non-zero temperature which we have to employ in our numerics, the spectrum shown in the plots has a finite width and the locations of the zero modes are approximately at the Fermi momentum (4.3). To investigate this point further and to confirm that we are indeed dealing with a genuine Fermi surface, in the following we compute the quasi-particle weight Z as a function of λ and g , as well as the effective mass and the lifetime at the Fermi surface as a function of ω , T , g and λ .

4.2.1 Quasi-particle residue

In order to scrutinize the Fermi surfaces, in this subsection we determine the quasi-particle residue by linearizing the dispersion relation around the Fermi surface at $\omega \approx 0$,

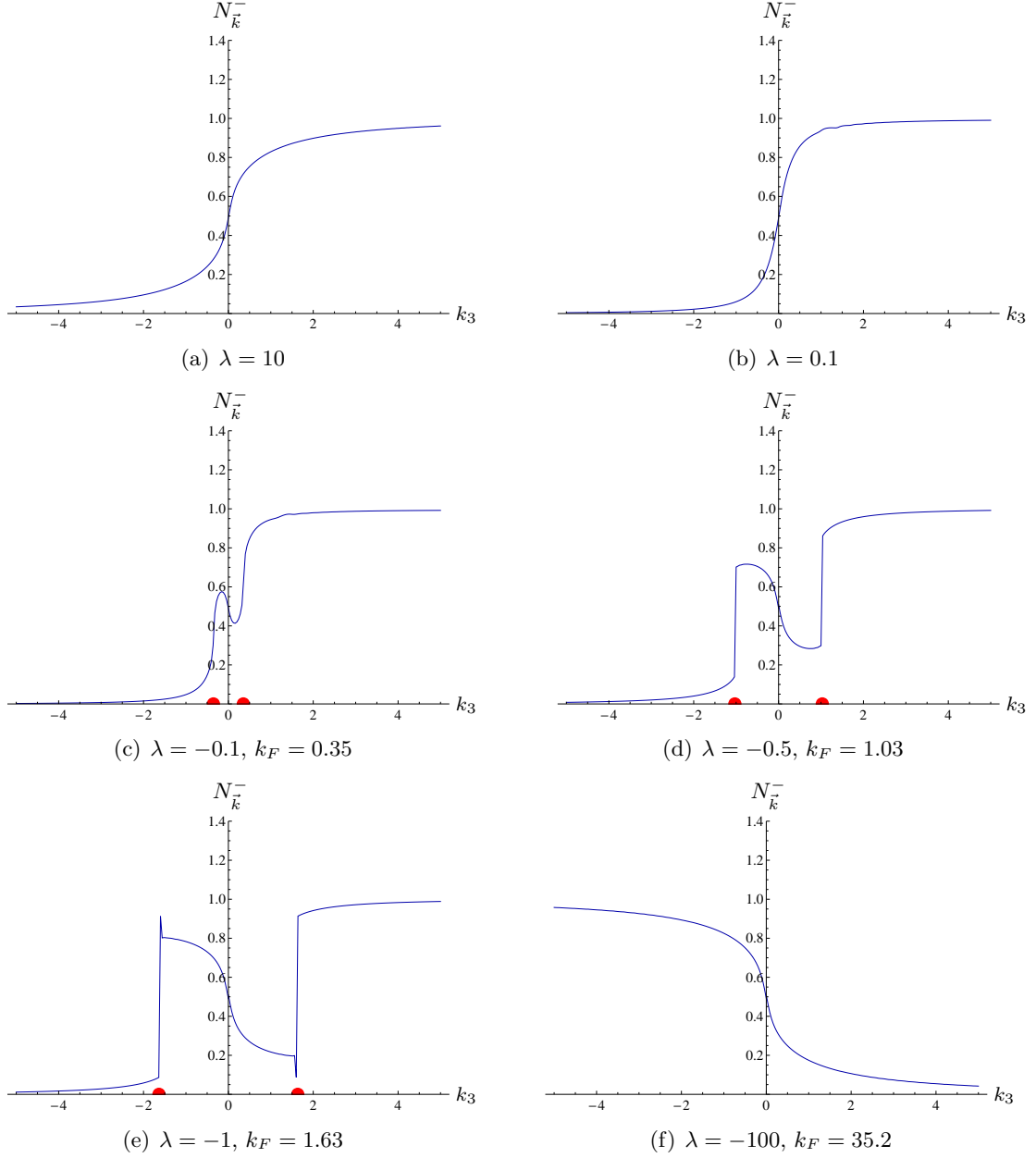


Figure 8: Momentum distribution of the minus component for $M = +1/4$, $T = 1/30$, $g = 1$. For $\lambda > 0$ it has a smooth kink-like behavior whereas for $\lambda < 0$ two Fermi surfaces develop. The red dots give the analytic value (4.3) of k_F . For small and negative λ , the Fermi surfaces are smeared out because of the non-zero value of the temperature.

$k \approx k_F$. For convenience, we choose the lower component of the Green's function which results in the spectral-weight function shown in figures 5(d)-(f). Furthermore, we employ the rotational symmetry to set $\vec{k} = (0, 0, k_3)$ with $k_3 > 0$, and thus $k_3 = k$ in the following. We then compute up to first order in derivatives

$$\begin{aligned} & - \operatorname{Re} (G_R^-)^{-1} \Big|_{k \approx k_F, \omega \approx 0} \\ & \approx \omega \left(1 - \partial_\omega \operatorname{Re} \Sigma^-(k_F, \omega) \Big|_{\omega=0} \right) - (k - k_F) \partial_k \left(\operatorname{Re} \Sigma^-(k, 0) - \frac{1}{\lambda} k^z \right) \Big|_{k=k_F}. \end{aligned} \quad (4.4)$$

Using (4.4) in the retarded Green's function, close to the Fermi surface we obtain the expression

$$G_R^-(k, \omega) = \frac{-Z}{\omega - Z(k - k_F) \partial_k \left(\operatorname{Re} \Sigma^-(k, 0) - \frac{1}{\lambda} k^z \right) \Big|_{k=k_F} - i Z \operatorname{Im} \Sigma^-(k, \omega)}, \quad (4.5)$$

where the wavefunction renormalization factor, or quasi-particle residue Z , is given by

$$Z = \frac{1}{1 - \partial_\omega \operatorname{Re} \Sigma^\pm(k_F, \omega) \Big|_{\omega=0}}. \quad (4.6)$$

The residue is equal for both spin components because the ω derivatives of $\operatorname{Re} \Sigma^\pm$ are equal at $\omega = 0$. Furthermore, note that Z depends on the coupling g explicitly due to the factor g in $\Sigma(p)$, but also implicitly via the dependence of k_F on g shown in (4.3). The dependence on λ is only through k_F . To be able to determine the quasi-particle residue, the real part of the self-energy has to be linear in $\omega = 0$, which is indeed the case, as shown in figure 9. The first derivative of the real part of the self-energy is shown for $T = 1/100$ in figure 10(a).

Let us illustrate the calculation of Z for a particular value of λ and g , namely $\lambda = -0.5$ and $g = 1$. For $T = 1/30$, the first derivative of the real part of the self-energy has the non-zero value $\partial_\omega \operatorname{Re} \Sigma^+(k_F, 0) \approx -0.8$ at $\omega = 0$. This leads to a finite quasi-particle residue of about $Z \approx 0.56$, which is precisely the height of the step in the momentum distribution at $k_3 = \pm k_F$, as shown in figure 11. Repeating this calculation for different values of λ , we have determined $Z(\lambda)$, which is shown in figure 12. Note that $0 \leq Z \leq 1$ as is required.

4.2.2 Effective mass

In the last subsection we have seen that at the Fermi surfaces there are quasi-particles which are strongly renormalized. Now, to further characterize the Fermi liquid we should also compute their effective mass. The latter is defined as the inverse of the slope of the quasi-particle dispersion around the Fermi surface, that is,

$$\omega(k) = \frac{1}{m_{\text{eff}}} (k - k_F) k_F \quad \text{near the Fermi surface}, \quad (4.7)$$

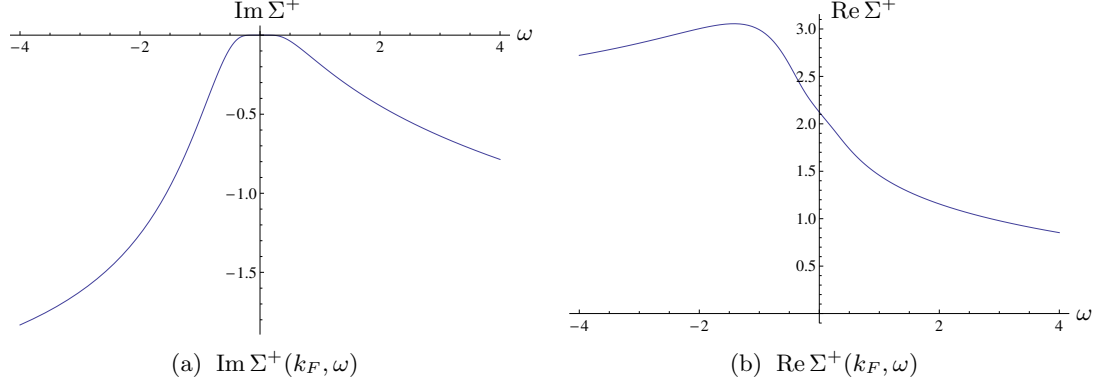


Figure 9: Imaginary and real part of the self-energy $\Sigma^+(p)$ evaluated at the Fermi momentum for $T = 1/100$, $g = 1$, $\lambda = -0.5$ and $M = 1/4$. The imaginary part is zero around $\omega = 0$ and the real part is linear around $\omega = 0$, which are both defining properties of a Fermi liquid.

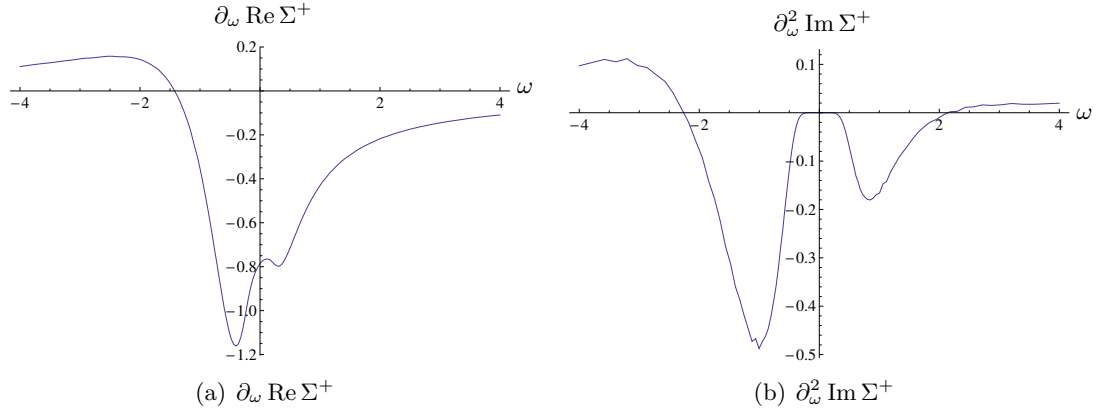


Figure 10: This figure shows the first derivative of $\text{Re } \Sigma^+(p)$ and the second derivative of $\text{Im } \Sigma^+(p)$, both at $\lambda = -0.5$, $T = 1/100$, $M = 1/4$ and $g = 1$, evaluated at the Fermi momentum.

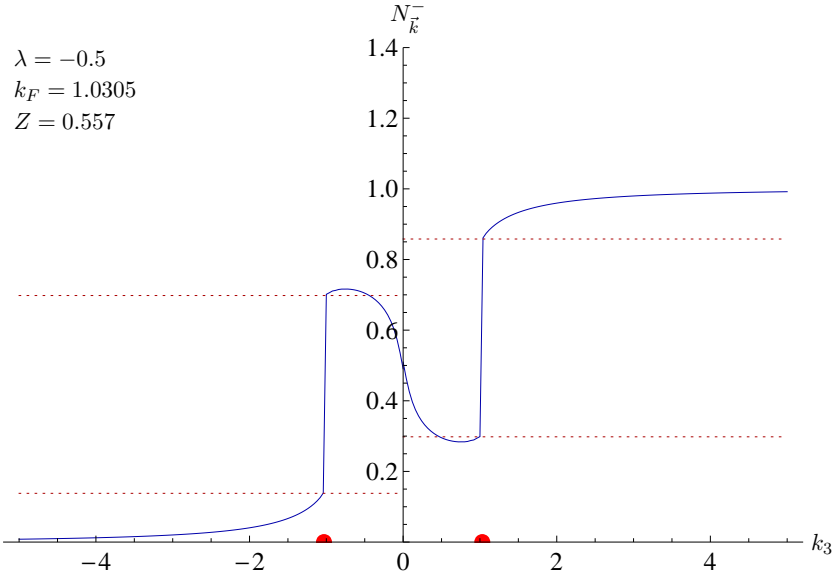


Figure 11: The momentum distribution for $M = 1/4$, $T = 1/30$ and $g = 1$. The quasi-particle residue is given by the distance between the dotted lines: the difference in height is the numerical value of Z , which is 0.56 for $\lambda = -0.5$. The red dots give the analytic value of $\pm k_F$.

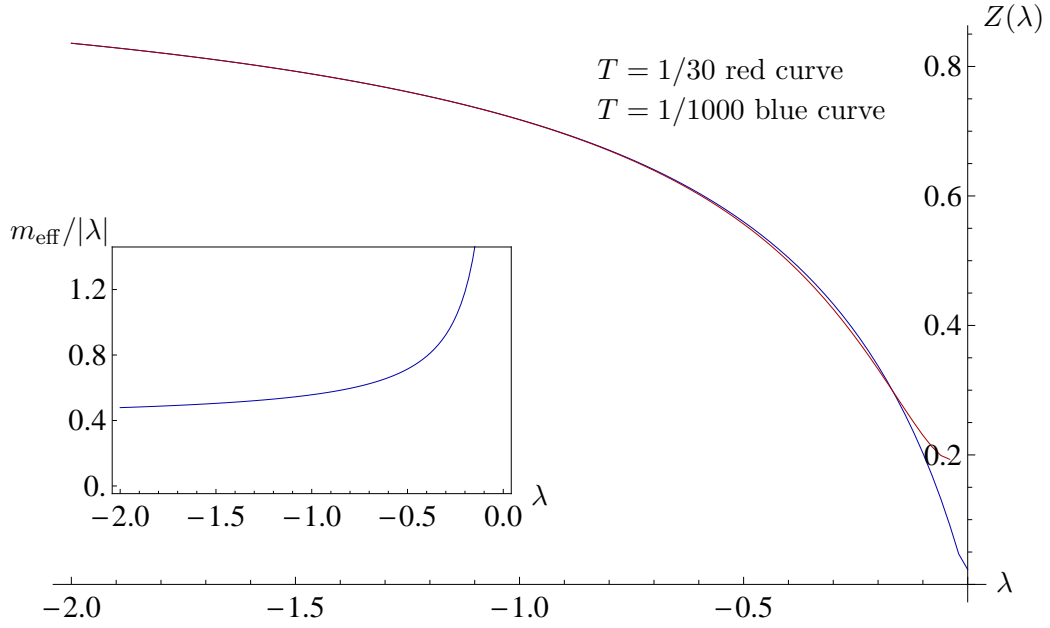


Figure 12: Wavefunction renormalization factor Z as a function of λ for $g = 1$, $M = 1/4$ and for different temperatures. Z is only defined for negative λ . The quasi-particle weight lies between one and zero, as expected. It approaches unity for very large and negative values of λ , i.e., very far away from the phase transition. For $\lambda \rightarrow 0^-$ it vanishes only strictly at $T = 0$. Due to non-zero temperature effects, the quasi-particle residue has a non-zero minimum at $\lambda = 0$, which is smaller for the lower temperature curve. The inset shows a plot of $m_{\text{eff}}/|\lambda|$ as a function of λ for $M = 1/4$.

where we are again employing $k_3 > 0$ and thus $k_3 = k$. Therefore, in our present situation the retarded Green's function near the Fermi momentum can be written as

$$G_R^-(k, \omega) = \frac{-Z}{\omega - \frac{1}{m_{\text{eff}}}(k - k_F)k_F - iZ \text{Im} \Sigma^-(k, \omega)} , \quad (4.8)$$

and by comparing with the explicit form given in (4.5) we conclude that

$$\frac{1}{m_{\text{eff}}} = \frac{Z}{k_F} \partial_k \left(\text{Re} \Sigma^-(k, 0) - \frac{1}{\lambda} k^z \right) \Big|_{k=k_F} . \quad (4.9)$$

At zero temperature, we can calculate the effective mass analytically. In particular, for $z = 2$ we obtain from the exact result given in equation (3.14) that

$$\text{Re} \Sigma^-(k, 0) = g c_1 k^{2M} , \quad \partial_k \text{Re} \Sigma^-(k, 0) \Big|_{k=k_F} = -\frac{2M}{\lambda} k_F , \quad (4.10)$$

where we eliminated $g c_1$ in favor of k_F . Therefore, the effective mass for $z = 2$ at zero temperature reads

$$m_{\text{eff}} = -\frac{\lambda}{2Z(\lambda)(M+1)} . \quad (4.11)$$

The effective mass is shown in the inset in figure 12.

4.2.3 Quasi-particle decay rate

We now turn to the lifetime of the quasi-particles. At zero temperature, from the analytic result (3.14) we infer that the imaginary part of the self-energy vanishes at $\omega = 0$, $k = k_F$. In figure 9, which shows the imaginary part of the self-energy, we see that this behavior is confirmed numerically. This implies that the spectral function at the Fermi surface is a δ -function at $T = 0$, which is a well-known and defining property of a Fermi liquid. As a consequence, the momentum distribution shows a discontinuity at the Fermi surfaces. Now, the inverse of the imaginary part of the self-energy $\text{Im} \Sigma^\pm$ has SI units of time, and at $\omega = 0$, we associate it to a lifetime for the quasi-particles⁴

$$\tau_{\vec{k}} = \frac{1}{-2Z \text{Im} \Sigma^\pm(\vec{k}, 0)} . \quad (4.12)$$

The lifetime τ_{k_F} is infinite precisely at the Fermi momentum. We can also define a frequency-dependent decay rate at the Fermi momentum as

$$\Gamma(\omega) = -2Z \text{Im} \Sigma^\pm(k_F, \omega) . \quad (4.13)$$

At the Fermi surface $\omega = 0$, the decay rate Γ is the inverse of the quasi-particle lifetime, therefore it vanishes for $\omega \rightarrow 0$. This behavior can also be observed in the spectra above.

⁴Here, we ignore ω -dependent corrections to the lifetime, as the dominant contribution comes from the behavior at $\omega = 0$.

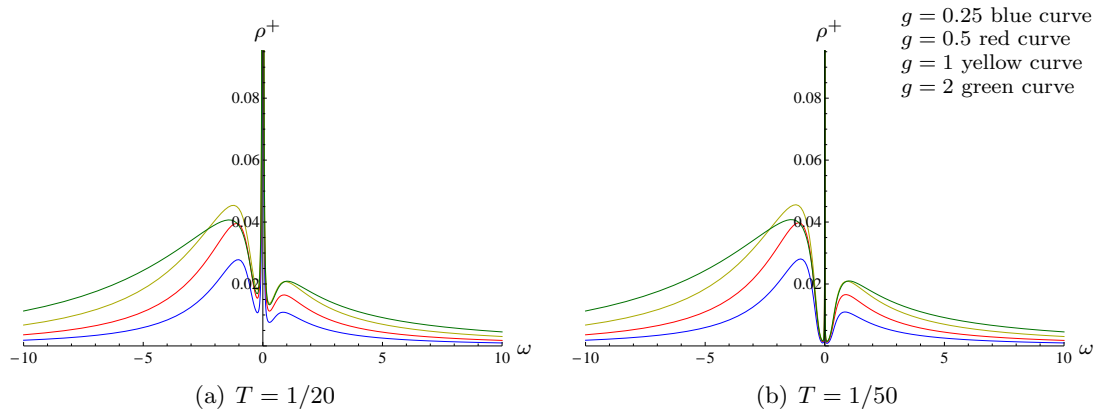


Figure 13: The spectral function for the spin-up component at the Fermi momentum as a function of ω . In both plots parameters are chosen as $g\lambda = -0.5$. This figure confirms numerically two important points that were made in the text. Firstly, for the higher temperature (a), a sharp peak is just visible at $\omega = 0$. For the lower temperature (b), this sharp peak is still there but is so narrow that it can barely be discerned. This indeed indicates that, at $T = 0$, there is a δ -function behavior at the Fermi surface, and that this δ -function peak becomes broader as temperature increases. Secondly, close to the Fermi surface, that is, for small ω , the form of the spectrum depends only on the product $g\lambda$. Indeed, all four curves for different g but equal $g\lambda$ overlap for small ω . This overlap is better for lower T . This shows that the scale of the Fermi momentum is indeed set by the combination $g\lambda$ and that the spectra are close to the Fermi surface indeed governed by k_F from (4.3) as we expect.

In particular, when T and ω are small but non-zero, the δ -function is broadened, which is indeed visible numerically, as shown in figure 13. For completeness, we note that the retarded Green's function given in (4.5) can now be written as

$$G_R^-(k, \omega) = \frac{-Z}{\omega - \frac{1}{m_{\text{eff}}}(k - k_F)k_F + \frac{i}{2\tau_k}}. \quad (4.14)$$

A broadening of the δ -function implies a non-zero decay rate for the quasi-particles. For a normal Fermi liquid, the imaginary part of the self-energy at the Fermi momentum vanishes at zero temperature and small frequency as ω^2 , which stems from Pauli blocking for scattering near the Fermi surface. For small non-zero temperatures, thermal averaging over ω^2 yields a T^2 dependence of the decay rate.

Let us compare this with our numerical results. We observe that the decay rate vanishes much faster than in a conventional Fermi liquid. In figure 10, we show the first derivative of $\text{Re}\Sigma^+(p)$ and the second derivative of $\text{Im}\Sigma^+(p)$ as a function of ω at $k = k_F$. The former is non-zero at $\omega = 0$, which indicates a non-zero value of Z as explained above. The latter is zero at $\omega = 0$, which shows that the quasi-particle decay rate vanishes faster than ω^2 . The behavior shown in figure 10 is in agreement with the existence of Fermi surfaces at $k = k_F$, however the decay rate is more suppressed than the typical ω^2 behavior of a conventional Fermi liquid. This indicates that, next to Pauli blocking, there is another mechanism at hand which strongly suppresses interactions

around the Fermi surface for low temperatures T and small frequencies ω . This appears to be a general feature of holographic Fermi liquids and can be understood better by performing a WKB analysis of the bulk Dirac equation.

4.3 Second-order WKB analysis around the Fermi surface

In section 4.2.3 we have seen that the fermionic quasi-particles are rather stable since their lifetime at the Fermi surface diverges faster than the power law expected from Fermi-liquid theory. To gain more insight into this property, we would like to obtain an analytic expression for the self-energy at the Fermi surface, that is, we want to know the form of the self-energy for both $\omega \neq 0$ and $k = k_F \neq 0$. As stated before, it is not possible to solve the differential equation for ξ exactly for generic non-zero values of ω and k , however, we can deploy the WKB approximation to study the fermionic excitations around the Fermi surfaces analytically, as was done for instance in [25, 26].

In principle, the WKB calculation gives the exact result for the retarded Green's function in the limit $k \rightarrow k_F$ and $\omega \ll 1$. Indeed, as we will see, it yields very elegantly the functional form of the self-energy at the Fermi surface, from which we can derive the behavior of the quasi-particle lifetime at $k = k_F$. The behavior turns out to be exponential, i.e.,

$$\text{Im } \Sigma \sim e^{-\alpha \frac{k^2}{\omega}}, \quad (4.15)$$

with α a constant to be determined. This result is consistent with the exponential behavior found in [25, 26]. Furthermore, we also find, to first approximation, an analytical result for the wavefunction renormalization Z , which is consistent with the strong quasi-particle renormalization found numerically. Notably, to obtain the latter result, we have to use the second-order WKB approximation.

However, before presenting the explicit result for the self-energy of our second-order WKB calculation, we have to mention the following. Obtaining the exact prefactors of the various terms in the final result, e.g. the constant α in (4.15), turns out to be non-trivial. The reason is that several integrals in the computation cannot be performed analytically (see appendix E for details) but have to be evaluated numerically. Alternatively, we could make an approximation to the integrals which, however, turns out to be similarly difficult. Our approach here is to deploy the WKB approximation to gain an understanding of the functional form of the quasi-particle lifetime at the Fermi surfaces. Therefore, the result that we present in equation (4.21) is obtained using the latter method, and so the prefactors, including α , are not exact. At the end of appendix E, we show some numerical evidence for the conjectured value of the constant α .

Let us outline the basic steps of our WKB calculation and refer to appendix E for further details. We determine the Green's function at *zero temperature*. The expansion parameter that plays the role of $1/\hbar$ in the conventional WKB approximation, and that we take to be large in order to satisfy the WKB conditions, is the rescaled momentum

(see appendix C.1)⁵

$$\bar{k} = \frac{k}{\sqrt{\omega}} \gg 1. \quad (4.16)$$

To simplify the calculation we also assume that both ω and \bar{k} are positive, and we introduce the rescaled variable $x = r/\sqrt{\omega}$. Choosing the up and down components u_{\pm} and d_{\pm} , respectively, of ψ_{\pm} introduced in (2.6) as

$$\psi_{\pm} = \frac{1}{\sqrt{r^5 V(r)}} \begin{pmatrix} u_{\pm} \\ d_{\pm} \end{pmatrix}, \quad (4.17)$$

we define new fluctuation fields y_{\pm} in terms of the original fields u_{\pm} as

$$y_{\pm} = \frac{\sqrt{x\bar{k} \pm 1}}{x^{3/2}} u_{\pm}. \quad (4.18)$$

Then, the once-iterated Dirac equation can be written as a Schrödinger-like equation

$$\frac{1}{\bar{k}^2} \frac{d^2}{dx^2} y_{\pm}(x) - V_{\pm}(x) y_{\pm}(x) = 0, \quad (4.19)$$

with the effective Schrödinger potential at $T = 0$ of the form

$$V_{\pm}(x) = \frac{1}{x^4} - \frac{1}{\bar{k}^2 x^6} + \frac{(\frac{1}{2} \pm M)(\frac{3}{2} \pm M)}{\bar{k}^2 x^2} + \frac{3}{4} \frac{1}{(x\bar{k} \pm 1)^2} - \frac{\frac{3}{2} \pm M}{x\bar{k}(x\bar{k} \pm 1)}. \quad (4.20)$$

The characteristic behavior of the Schrödinger potentials for the chiral components is illustrated in figure 14. In particular, V_+ has a single classical turning point at x_{tp} and V_- has a pole and a turning point. For $\bar{k} \gg 1$, both effective potentials have a very high potential barrier close to $x \rightarrow 0$. For simplicity we present here the calculation only for the plus-component, the result for the minus-component is easily obtained using the symmetry explained in appendix C. Also, we only perform the calculation for the case $M > 0$ so that the single turning point of V_+ at x_{tp} divides the entire range into two regions, a classically allowed region where $V < 0$ and a classically disallowed region where $V > 0$.

The strategy to determine the WKB solution is as follows. We first incorporate the in-falling near-horizon boundary condition into the WKB wavefunction in the classically allowed region $x < x_{\text{tp}}$. Then, we apply the usual linearized connection formulae at $x = x_{\text{tp}}$ to continue the solution into the classically disallowed region $x > x_{\text{tp}}$. Next, we match this solution onto the exact near-boundary solution to fix the remaining integration constants. Finally, from the asymptotic expansion of the solution we read off the coefficients A_{\pm} (see appendix B.2) that determine G_R via $\xi = iA_-/A_+$. The details of

⁵We will specialize to the case of interest $z = 2$ in this section, although the calculation can easily be generalized to arbitrary $z > 1$.

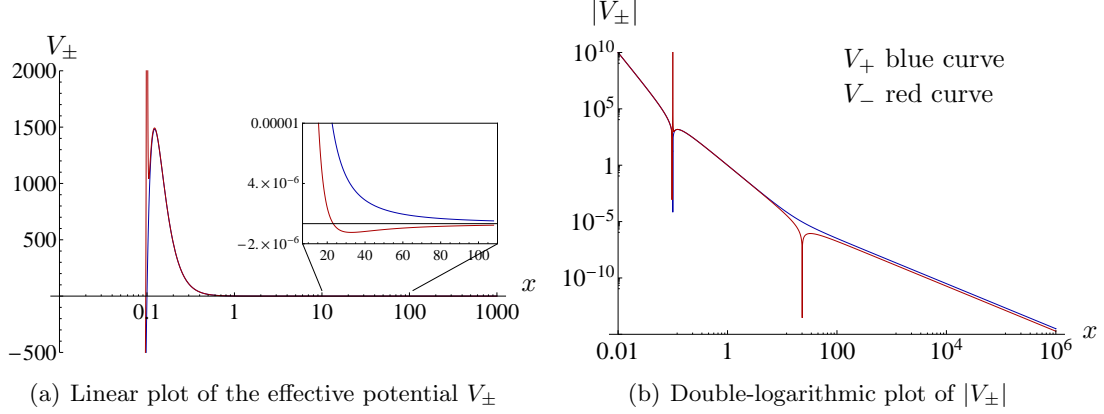


Figure 14: Plots of the effective potentials V_+ and V_- (blue and red line) as function of $x = r/\sqrt{\omega}$ at $T = 0$ for $\bar{k} = 10$ and $M = 1/4$. In the near-horizon region $0 < x < x_{\text{tp}}$, both effective potentials are negative. The potential V_+ has a turning point at x_{tp} , which can be seen as the first dip from the left in (b). V_- has a pole located approximately at the position of the classical turning point of V_+ . V_+ is always positive for $x > x_{\text{tp}}$, while V_- has a different turning point at an intermediate value of x . This point can be seen in the inset in (a) and as the second dip from the left in (b).

this calculation are explained in appendix E, and the final result for the WKB self-energy reads

$$\Sigma^+(\vec{k}, \omega) = g c_1 k^{2M} \frac{\gamma_+}{\epsilon_+} \frac{1 - \frac{i}{2} \frac{\gamma_-}{\gamma_+} e^{-i\pi M} e^{-2\frac{k^2}{\omega}}}{1 + \frac{i}{2} \frac{\epsilon_-}{\epsilon_+} e^{i\pi M} e^{-2\frac{k^2}{\omega}}}, \quad (4.21)$$

where we employed again $\vec{k} = (0, 0, k_3)$ with $k_3 > 0$, so that the corresponding Green's function in the limit $\bar{k} \gg 1$ reads

$$G_R^+(\vec{k}, \omega) = - \left(\omega - \frac{1}{\lambda} k^z - g c_1 k^{2M} \frac{\gamma_+}{\epsilon_+} \frac{1 - \frac{i}{2} \frac{\gamma_-}{\gamma_+} e^{-i\pi M} e^{-2\frac{k^2}{\omega}}}{1 + \frac{i}{2} \frac{\epsilon_-}{\epsilon_+} e^{i\pi M} e^{-2\frac{k^2}{\omega}}} \right)^{-1}. \quad (4.22)$$

The constant c_1 was given in (3.2), and we defined

$$\epsilon_{\pm} = 1 \pm \frac{M}{2\bar{k}^2} \log \bar{k} \mp \frac{1}{4\bar{k}^2}, \quad \gamma_{\pm} = 1 \pm \frac{M}{2\bar{k}^2} \log \bar{k} \mp \frac{3}{4\bar{k}^2}. \quad (4.23)$$

As mentioned before, the result (4.21) comes from performing an expansion *within* the framework of the WKB calculation. In this case, we find $\alpha = 2$ for the aforementioned coefficient in (4.15), which we can improve by taking into account more terms in the expansion. In particular, as explained in more detail in appendix E, the correct prefactor in the exponent seems to be approximately $\alpha \approx \pi/2$.

After having obtained our result, let us now study the Green's function (4.22) in more detail. As a first consistency check, we see that in the limit $\omega \rightarrow 0$, which corresponds to

$\bar{k} \equiv k/\sqrt{\omega} \rightarrow \infty$, we find the same locus for the Fermi surface at $k = k_F$ as in equation (4.3). Furthermore, the decay rate of the quasi-particle excitations around the Fermi surface is calculated using (4.12) and employing the imaginary part of the self-energy at $k = k_F$, which reads

$$\text{Im } \Sigma^+(\vec{k}_F, \omega) = \frac{1}{2} g c_1 k_F^{2M} \sin\left(\pi\left(M - \frac{1}{2}\right)\right) \frac{\gamma_+}{\epsilon_+} \left(\frac{\gamma_-}{\gamma_+} + \frac{\epsilon_-}{\epsilon_+}\right) e^{-2\frac{k_F^2}{\omega}}, \quad (4.24)$$

where k_F is given by (4.3). We note that γ_{\pm} and ϵ_{\pm} become unity in the strict Fermi-surface limit, and the result thus shows an exponentially suppressed decay rate with additional ω/k_F^2 corrections. This result demonstrates the extremely stable nature of the quasi-particle excitations on the Fermi surface and is consistent with previous WKB calculations that similarly find exponentially narrow decay rates [25, 26]. A comparison with the numerical results is made in figure 15.

Another analytic result that we extract from (4.21) is the wavefunction renormalization (4.6) on the Fermi surface. This follows from the real part of the self-energy which reads

$$\text{Re } \Sigma^+(\vec{k}, \omega) = g c_1 k^{2M} \frac{\gamma_+}{\epsilon_+}, \quad (4.25)$$

up to exponentially small terms multiplied by $\exp(-\alpha k^2/\omega)$ that will drop out in the limit $\omega \rightarrow 0$ after differentiation. Combining (4.25) with (4.3) and (4.6), we arrive at an analytic result for the wavefunction renormalization factor of the form

$$Z = \frac{2\lambda}{2\lambda - 1}, \quad (4.26)$$

which for $\lambda < 0$ satisfies $0 \leq Z \leq 1$ as required. We note that, to obtain this non-trivial result different from one, we have to go beyond the leading order in the WKB approximation, since otherwise the derivative of the self-energy in (4.21) would vanish due to the exponential suppression. Interestingly, because all factors M in (4.23) are multiplied by ω , these drop out in the limit $\omega \rightarrow 0$ so that the analytic result (4.26) is independent of M . This is in contrast to the numerical results which do depend on the value of M . The analytic result (4.26) is therefore expected to be the $M = 0$ part of a larger expression, that one may obtain using a more sophisticated WKB calculation. This presumption is supported by figure 16, which shows that both results are consistent in the case $M = 0$. However, the aim of the WKB calculation was merely to support our numerical results, and finding perfect agreement is beyond the scope of this work.

4.4 Phase diagram

The results obtained in the sections above, in particular the plots in figure 8, can be explained by the existence of a quantum phase transition at $\lambda = 0$ and $g \geq 0$. Firstly, we can understand why there must be different phases. Consider again the Green's function (2.11) and recall that λ can change sign whereas g is required to be positive (in the range

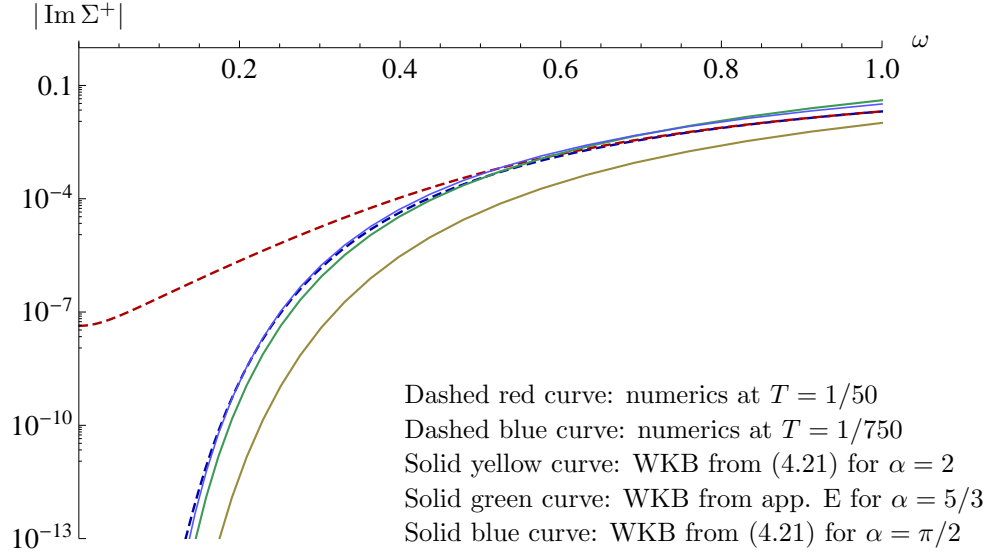


Figure 15: Comparison of WKB and numerical results for $|\text{Im } \Sigma^+|$ at $k = k_F$ with $\lambda = -1$ yielding $k_F \approx 1.64$, and $M = 1/4$ as a function of ω . The dashed curves are numerical results, the solid curves are second-order WKB results at $T = 0$. Note that the WKB result (4.21) for $\alpha = \pi/2$ agrees very well with the low temperature numerical result.

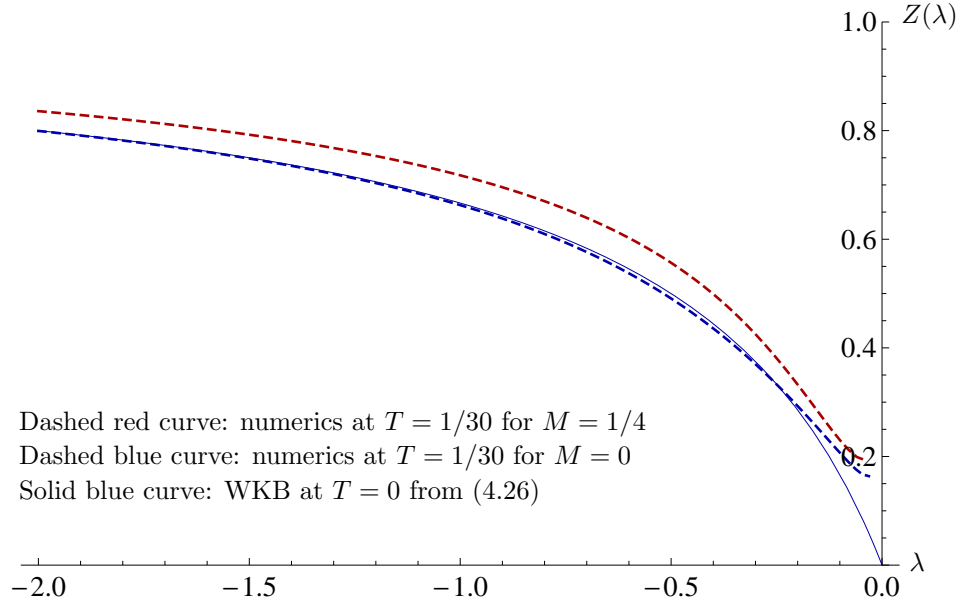


Figure 16: Comparison of WKB and numerical results for the quasi-particle residue Z from (4.6) as a function of λ . The dashed curves are numerical results, the solid curve is the second-order WKB result at $T = 0$. This figure shows that the WKB result for Z resembles the $M = 0$ result, indicating that the analytic expression may be the $M = 0$ part of a more general result.

$-1/2 < M < 1/2$) in order to satisfy the Kramers-Kronig relation. Furthermore, the kinetic term scales as k^z , but the self-energy is proportional to k^{2M} which is a lower power of k . As a consequence, there will always be a small but non-zero \vec{k} interval at low values of the momentum in which the self-energy term dominates. When there is a difference in the relative sign, there can be a cancellation between the kinetic and the interaction term in the Green's function at low momenta. This is what brings the Fermi surfaces into existence. When there is no relative sign difference, the self-energy contribution adds up to the kinetic term, which results in a different behavior of $N_{\vec{k}}^{\pm}$ and therefore in a different phase. Thus, the crucial property is the possibility of a relative sign change between the kinetic and interaction term which is dominant at low momenta.

We sketch the phase diagram of our system in figure 17, and relate it to our results for the momentum distribution at various values of λ and T .

- For $\lambda < 0$ and temperatures T small compared to the Fermi energy $\epsilon_F = k_F^z/|\lambda|$, we have a phase with two Fermi surfaces: one for the particles and one for the holes, which can be seen as two sharp discontinuities in the momentum distribution at $k_3 = \pm|k_F|$. The system is formally only a Fermi liquid at exactly zero temperature. Numerically we can obtain only non-zero temperature results, and the effect of temperature is to smear out the discontinuities over an interval around $\pm|k_F|$. However, when T/ϵ_F is small (outside the critical region) we still see the Fermi-liquid-like or degenerate behavior. Indeed in figure 8 for small $\lambda < 0$ the jumps are not infinitely sharp but still very much located around a single value of \vec{k} .
- When increasing temperature, in the region $T \sim \epsilon_F$ there is a crossover between the Fermi-liquid phase and the quantum critical phase, indicated by the dashed lines in figure 17. To calculate the crossover temperature as a function of λ we realize that the crossover occurs when $T \approx \epsilon_F$, thus at

$$T \approx \epsilon_F = \frac{1}{|\lambda|} k_F^z = |\lambda|^{\frac{2M}{z-2M}} = g c_1 k_F^{2M}. \quad (4.27)$$

In the case $M = 1/4$, $z = 2$, the crossover is at $T \sim |\lambda|^{1/3}$.

- When $T \gg \epsilon_F$ the Fermi surfaces will be smeared out to an extent that they are difficult to identify, as it can be seen in figure 18(a). Then the system is in the quantum critical phase. As $\lambda \rightarrow 0^-$, ϵ_F decreases, and as a consequence the Fermi surfaces develop into smooth bumps for very small λ . Numerically, we will always observe temperature effects for small values of $|\lambda|$.
- At $\lambda = 0$ and $T = 0$, there is a quantum phase transition from a phase with two Fermi surfaces to a phase with no Fermi surfaces. At $\lambda = 0$ and $T > 0$, the system is in the quantum critical phase. Interestingly, there is a qualitative difference between the quantum critical $\lambda < 0$ and $\lambda > 0$ phases, see figure 18(b) for an impression of the behavior of $N_{\vec{k}}^{\pm}$ for very small λ .

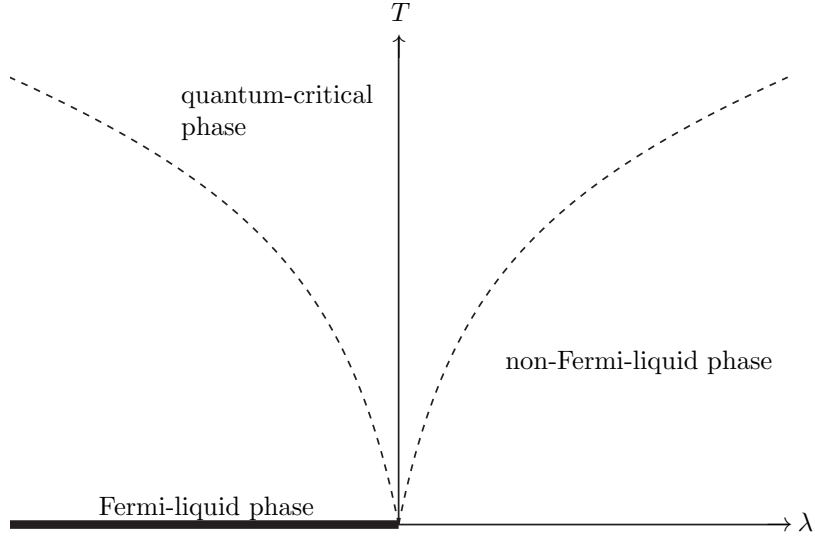


Figure 17: Sketch of the phase diagram in the case $-1/2 < M < 1/2$ and $g = 1$, showing the non-Fermi-liquid phase for low T and positive λ , the Fermi-liquid phase for low T and negative λ , and two qualitatively different quantum-critical phases for high temperatures and λ positive and negative. The system is strictly only a Fermi liquid at $T = 0$ (thick line), but has Fermi-liquid-like behavior for low temperatures. The dashed curves correspond to the crossovers to the quantum-critical regimes. There is a quantum phase transition at $\lambda = 0, T = 0$.

- In the case $\lambda > 0$, the momentum distribution has a kink-like behavior and is centered around $k = 0$. For $T \ll \epsilon_F$, the width of the kink is determined by $k_F \equiv (g\lambda c_1)^{1/(z-2M)}$, whereas for $T \gg \epsilon_F$ it is determined by T , see figure 18(c).

4.5 Fermi surfaces and the quantum phase transition for $\frac{1}{2} < |M| < \frac{z}{2}$

In the previous sections, we have studied the system in the particular range $0 < M < 1/2$ and we have demonstrated the appearance of Fermi surfaces for $\lambda < 0$. However, we obtain a similar behavior also for positive values of λ if we consider $1/2 < M < z/2$ and $-z/2 < M < -1/2$.

Let us demonstrate this for $1/2 < M < z/2$. In this case, the relation between the self-energy $\Sigma(p)$ and the bulk quantity $\xi(r, p)$ is given by equation (2.14). However, the expression (C.24) for the Green's function is unchanged because this form is determined solely by the symmetry properties of $\xi(r, p)$, which are also satisfied by the counter-term in (2.14). In particular, the $\omega \rightarrow 0$ limit of the dispersion relation is again given by equation (4.2) which we recall for convenience,

$$\frac{1}{\lambda} k^z + g k^{2M} c_1 = 0. \quad (4.28)$$

Now, we can find a non-trivial, real solution to this equation in the range $1/2 > M > z/2$ for *positive* λ because the constant c_1 given by (3.2) is negative in this range. Note that the same conclusion applies for $-z/2 < M < -1/2$. Thus, we now obtain Fermi

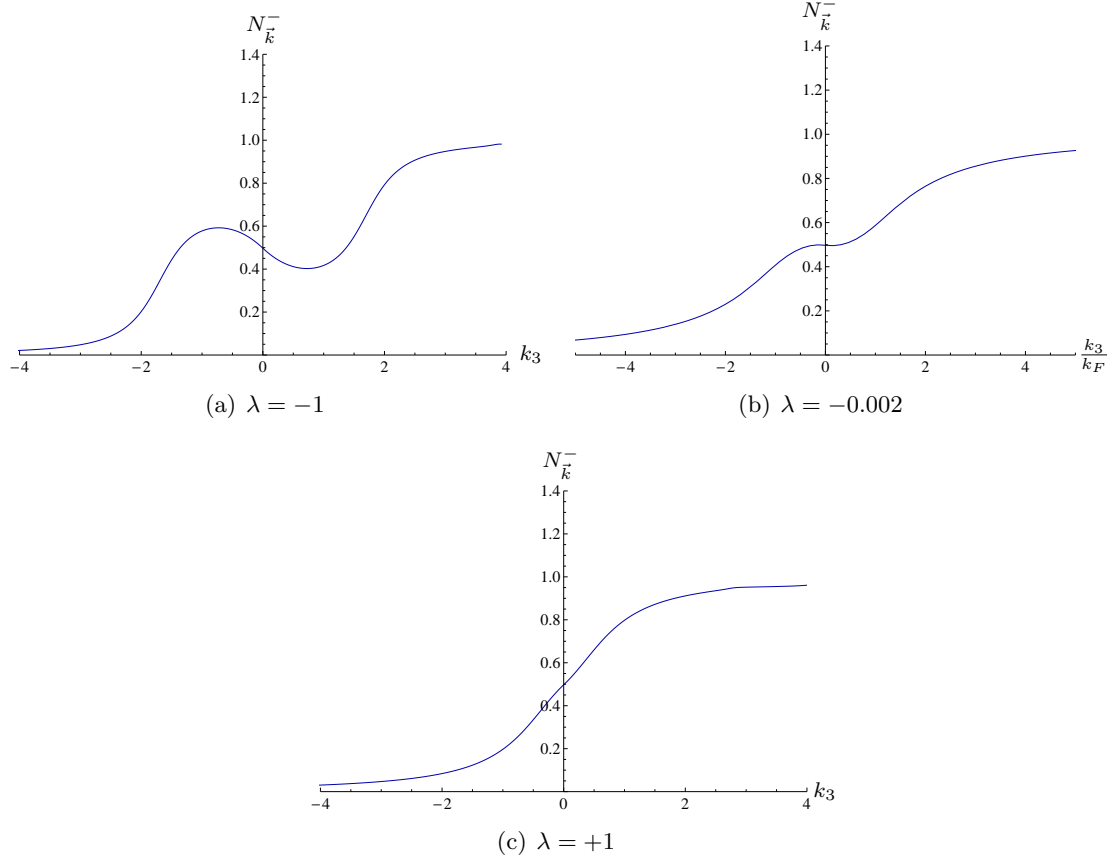


Figure 18: Momentum distribution in the quantum critical region for $T = 5$, $g = 1$, $M = 1/4$ and various values of λ . For $\lambda < 0$, the momentum distribution always shows two extrema at non-zero k_3 , which develop into Fermi surfaces when the temperature is lowered. This is shown in (a). The behavior for small negative λ is illustrated in figure (b) with $k_F \approx 0.026$. Note that because k_F is relatively small, we plotted $N_{\vec{k}}^-$ here as a function of k_3/k_F for clarity. For $\lambda > 0$ the momentum distribution is very broad as can be seen in (c).

surfaces for positive λ in these ranges for M . Furthermore, we also find a quantum phase transition at $\lambda = 0$ below which the aforementioned Fermi surfaces disappear.

5 Conclusion and discussion

In our recent work [14] we proposed the construction of a holographic fermionic retarded Green's function modified to describe single-particle correlations, using the AdS/CMT correspondence. Two aspects of this prescription that make it particularly interesting for condensed-matter applications, are the following. Firstly, the retarded Green's function satisfies the zeroth frequency sum rule and the Kramers-Kronig relation both at zero and also at non-zero temperatures. We have checked that this is the case analytically where possible, see appendix D, and otherwise numerically. This shows that the correlation function is a genuine single-particle correlation function that can be compared to experimental ARPES data. Secondly, it can describe quantum critical points with relativistic as well as non-relativistic dynamical exponent $z > 1$, as a consequence of the (approximate) Lifshitz isometry of the used bulk spacetime.

In this paper, we have investigated the physics arising from this modified holographic prescription. The resulting spectral-weight function and dispersion relations that we presented in section 3, show that we are describing a particle-hole symmetric, chiral boundary system that is a strongly interacting Weyl semimetal. Weyl semimetals have recently received considerable interest in the condensed-matter community because of their fascinating and unusual properties that result from the topological nature of the band structure [27, 28, 29, 30, 31, 32]. In order to investigate the band structure, the Weyl semimetals are usually treated as free or weakly interacting systems. Using our holographic prescription of Weyl semimetals, we are now also able to explore these systems when they are strongly interacting or critical. In particular, the inclusion of a holographic self-energy leads to the existence of several different phases including a Fermi-liquid phase and a non-Fermi-liquid phase, separated by a quantum phase transition, as we have shown in section 4.

There are various interesting directions in which our prescription can be extended. Firstly, our current prescription is intended to describe single-particle correlation functions, which appear often in condensed matter. However, in experiments one also encounters two-particle correlation functions. Important examples include the current-current correlation function that determines the electrical conductivity, σ , and the heat current-heat current correlation function which is related to the thermal conductivity κ . The electrical conductivity has been investigated in a holographic context before, starting with [33], but it would be interesting to also consider this particular function using our holographic prescription. In particular, since our present discussion is at zero doping, that is, it contains an equal amount of particles and holes, we expect the corresponding conductivity in the boundary system to remain finite in the $\omega \rightarrow 0$ limit even without impurities. Additionally, in Fermi-liquid theory the thermal conductivity κ scales as $1/T$, whereas the electrical conductivity σ scales as $1/T^2$. Their ratio is therefore proportional to temperature, i.e., $\kappa/\sigma \sim T$, which is the famous Wiedemann-Franz law. It

would be interesting to investigate whether this law also holds here using the holographic results for electrical and thermal conductivity.

As stressed before, the various phases and the corresponding quantum phase transition described in section 4 are realized at zero doping. The scale that governs the quantum phase transition is an appropriate combination of the parameters g and λ appearing in our prescription. As a generalization of our holographic description of a Weyl semimetal we will extend the prescription to non-zero doping in future work. This is interesting for condensed-matter purposes as it corresponds for instance to turning on a gate voltage in the material. Holographically this can be done by adding another U(1) gauge field to the bulk spacetime under which the fermions are charged, and by considering charged Lifshitz black branes [34]. It will be interesting to investigate the consequences for the phase diagram described in section 4.

A topic often discussed in the AdS/CMT literature is the well-known Luttinger theorem from Fermi-liquid theory, which states that the total particle density in a system is proportional to the total volume in momentum space enclosed by its Fermi surfaces. In our case, this can be checked as follows. The system we describe is at zero doping, so the total number density of the particles and holes together is zero. The Fermi surfaces for particles and holes separately enclose a non-zero volume, but the difference in volume enclosed by the Fermi surfaces is zero. Therefore, the Luttinger theorem is trivially satisfied in our case.

Finally, whereas our current prescription takes into account only the leading-order contributions in the $1/N$ expansion, it would be interesting to also consider quantum or $\mathcal{O}(1/N)$ corrections. In particular, it is sometimes proposed [35] that the typical feature of holographic Fermi liquids, an exponentially suppressed quasi-particle decay rate at the Fermi surface, may be a large N remnant, i.e., a consequence of the fact that the leading order does not take into account all possible quasi-particle decay processes. Therefore, considering also $1/N$ contributions may resolve this problem and restore the conventional ω^2 power-law behavior, which is rather robust in Fermi liquids. An example of $1/N$ corrections that are ignored here, are the back-reaction of the massive fermions on the gravitational background in the bulk, but one may also think of other $1/N$ corrections.

Acknowledgments

It is a pleasure to thank R. Duine, C. Herzog, S. Hartnoll and S. Sachdev for discussions. This work was partially supported by the Netherlands Organization for Scientific Research (NWO) under the VICI grant 680-47-603.

A Conventions

A.1 A note on dimensions and units

In this paper, we work with natural units in which $\hbar = c = k_B = 1$, where k_B is Boltzmann's constant, and we use dimensionless coordinates r, t, \vec{x} . As a consequence, Newton's constant G_{d+1} as well as the metric are dimensionless. However, to convert the Lifshitz metric (2.1) to standard SI units, we introduce a length scale l and define the dimensionful expression

$$d\tilde{s}^2 \equiv l^2 ds^2 = \frac{l^2}{\tilde{r}^2} \frac{d\tilde{r}^2}{V^2(\tilde{r})} - V^2(\tilde{r}) \frac{\tilde{r}^{2z}}{l^{2z}} c^2 d\tilde{t}^2 + \frac{\tilde{r}^2}{l^2} d\vec{\tilde{x}}^2, \quad (\text{A.1})$$

where

$$\tilde{r} = l r, \quad \vec{\tilde{x}} = l \vec{x}, \quad \tilde{t} = \frac{l}{c} t, \quad (\text{A.2})$$

which has the correct units of meter and second, respectively. Note that l can be interpreted as the scale characterizing the size of the Lifshitz spacetime. The temperature of the black brane, obtained in (2.2), can be converted to have dimension of Kelvin by the rescaling

$$\tilde{T} = \frac{\hbar c}{k_B} \frac{T}{l}, \quad (\text{A.3})$$

and the mass of the fermions in SI-units takes the form

$$\tilde{M}^2 = \frac{\hbar c}{\tilde{G}_{d+1}} M^2 l^{d-3}. \quad (\text{A.4})$$

Here, \tilde{G}_{d+1} is Newton's constant which has dimension $\text{m}^d \cdot \text{s}^{-2} \cdot \text{kg}^{-1}$. To convert the dimensionless frequencies ω and momenta \vec{k} to SI units, one should rescale

$$\tilde{\omega} = \frac{c}{l} \omega, \quad \vec{\tilde{k}} = \frac{\vec{k}}{l}. \quad (\text{A.5})$$

Employing the results of this section, the reader can at any time convert to physical units. However, for ease of notation, we present our analysis and results in dimensionless coordinates.

A.2 Vielbeins and Dirac matrices

Here, we collect some formulae related to the spin connection and vielbeins, and state our choice of Dirac matrices. The vielbeins are denoted by e^a_μ with $\mu = 0, \dots, d$ and $a = 0, \dots, d$, and satisfy

$$g_{\mu\nu} = e^a_\mu \eta_{ab} e^b_\nu, \quad \eta_{ab} = \text{diag}(-1, +1, \dots, +1). \quad (\text{A.6})$$

We will furthermore employ the following notation and relations

$$e^a = e^a_\mu dx^\mu, \quad \Omega^a_{\underline{b}} = (\Omega_\mu)^a_{\underline{b}} dx^\mu, \quad 0 = de^a + \Omega^a_{\underline{b}} \wedge e^{\underline{b}}, \quad (\text{A.7})$$

where Ω is the torsion-free spin connection. Given the metric (2.1), we can read off

$$e^r = \frac{dr}{rV}, \quad e^t = r^z V dt, \quad e^i = r dx^i, \quad (\text{A.8})$$

where $\underline{a} = \{t, r, i\}$ with $i = 2, \dots, d$ labeling the spatial directions. Using the relation in (A.7) as well as the anti-symmetry of Ω , we determine its non-vanishing components as

$$\Omega_{\underline{r}}^t = rV \partial_r (r^z V) dt, \quad \Omega_{\underline{r}}^i = rV dx^i. \quad (\text{A.9})$$

For the spinors, we use the following notation

$$\bar{\Psi} = \Psi^\dagger \Gamma^t, \quad \mathcal{D}_\mu = \partial_\mu + \frac{1}{4} (\Omega_\mu)_{\underline{ab}} \Gamma^{\underline{ab}}, \quad \mathcal{D} = \Gamma^a e_{\underline{a}}^\mu \mathcal{D}_\mu. \quad (\text{A.10})$$

Here, $e_{\underline{a}}^\mu$ is the inverse transpose of $e_\mu^{\underline{a}}$ and the symbol $\Gamma^{\underline{ab}}$ is defined in terms of the Gamma matrices $\Gamma^{\underline{a}}$ as

$$\Gamma^{\underline{ab}} = \frac{1}{2} [\Gamma^{\underline{a}}, \Gamma^{\underline{b}}]. \quad (\text{A.11})$$

We also perform chiral projections using

$$\Psi_\pm \equiv \frac{1}{2} (1 \pm \Gamma^r) \Psi, \quad \Gamma^r \Psi_\pm = \pm \Psi_\pm. \quad (\text{A.12})$$

A convenient choice of Gamma matrices reads as follows [36]:

- For odd dimensions d , we can choose

$$\Gamma^r = \begin{pmatrix} \mathbb{1} & 0 \\ 0 & -\mathbb{1} \end{pmatrix}, \quad \Gamma^t = \begin{pmatrix} 0 & \gamma^0 \\ \gamma^0 & 0 \end{pmatrix}, \quad \Gamma^i = \begin{pmatrix} 0 & \gamma^i \\ \gamma^i & 0 \end{pmatrix}, \quad (\text{A.13})$$

where the γ -matrices are the d -dimensional Dirac matrices of the boundary theory. We then decompose Ψ into chiral components of the bulk spinor

$$\Psi = \begin{pmatrix} \psi_+ \\ \psi_- \end{pmatrix}, \quad (\text{A.14})$$

where the components ψ_\pm are chiral in the bulk, but they are Dirac spinors on the boundary.

- In the case when d is even, a natural choice is given by

$$\Gamma^r = \gamma^{d+1}, \quad \Gamma^t = \gamma^0, \quad \Gamma^i = \gamma^i, \quad (\text{A.15})$$

where γ^{d+1} is the analog of γ^5 in four dimensions and γ are the gamma matrices of the boundary theory. In this case Ψ_\pm correspond to operators with *definite chirality* on the boundary.

- Most of our interest lies in the case $d = 4$. Using (A.12), and with a slight abuse of notation, we write the four-component Dirac bulk spinor Ψ in terms of two-component spinors Ψ_+ and Ψ_- as

$$\Psi = \begin{pmatrix} \Psi_+ \\ \Psi_- \end{pmatrix} . \quad (\text{A.16})$$

The four-dimensional gamma matrices can be expressed in terms of $\sigma^a = (1, \vec{\sigma})$ and $\bar{\sigma}^a = (-1, \vec{\sigma})$ with σ^i being the Pauli matrices in the following way

$$\Gamma^a = \begin{pmatrix} 0 & \bar{\sigma}^a \\ \sigma^a & 0 \end{pmatrix} , \quad \underline{a} = \{\underline{t}, \underline{i}\} . \quad (\text{A.17})$$

B Details on the near-boundary asymptotics

B.1 Dirac equation

The Dirac equation for the fermion field Ψ can be obtained from varying the action (2.4) with respect to $\bar{\Psi}$, and imposing $\delta\Psi_+ = 0$ at $r = r_0$ (see [14] for more details). We find the usual expression $(\not{D} - M)\Psi = 0$, which can be written as follows

$$\left[rV \Gamma^r \partial_r + \frac{i}{r} \Gamma \cdot \tilde{k} + \frac{1}{2} \Gamma^r p_z(r) - M \right] \psi(r) = 0 . \quad (\text{B.1})$$

Here we defined the function $p_z(r) = r^{1-z} \partial_r [r^z V] + (d-1)V$, and used the notation $\Gamma \cdot \tilde{k} = \Gamma^{\underline{t}} \tilde{\omega} + \Gamma^{\underline{i}} k_i$ together with the *generalized momenta*

$$\tilde{k} = (\tilde{\omega}, k_i) , \quad \tilde{\omega} = -\frac{\omega}{r^{z-1}V} . \quad (\text{B.2})$$

Next, we can simplify equation (B.1) by introducing

$$\psi(r) = \frac{1}{\sqrt{r^{d-1+z}V}} \phi(r) , \quad (\text{B.3})$$

which leads to

$$\left[rV \Gamma^r \partial_r + \frac{i}{r} \Gamma \cdot \tilde{k} - M \right] \phi(r) = 0 . \quad (\text{B.4})$$

In terms of the chiral components $\phi_{\pm}(r)$, the Dirac equation (B.4), both in even and odd dimensions d , then reads

$$\phi_{\pm}(r) = \mp \frac{i}{\tilde{k}^2} (\gamma \cdot \tilde{k}) \mathcal{A}(\mp M) \phi_{\mp}(r) , \quad \mathcal{A}(M) \equiv r(rV \partial_r - M) , \quad (\text{B.5})$$

where we employed the notation for the Γ matrices given in appendix A.2. Finally, we can also derive a *second-order* equation by applying $\mathcal{A}(\pm M)(\gamma \cdot \tilde{k})$ to both sides of the Dirac equation in (B.5), which leads to

$$\tilde{k}^2 \phi_{\pm}(r) = \mathcal{A}(\mp M) \mathcal{A}(\pm M) \phi_{\pm}(r) - r^2 V \frac{\partial_r \tilde{\omega}}{\tilde{k}^2} \gamma^0 (\gamma \cdot \tilde{k}) \mathcal{A}(\pm M) \phi_{\pm}(r) . \quad (\text{B.6})$$

Note that for the case of interest in this paper, that is $d = 4$, half of the components in equations (B.5) and (B.6) are trivial, because we can use the Weyl representation of the gamma matrices so that ϕ_{\pm} correspond to the upper and lower two components of ϕ .

B.2 Asymptotic solutions near the boundary

We now determine the form of the asymptotic solutions near the boundary at $r = \infty$. In contrast to the AdS case with $z = 1$, for $z \neq 1$, the asymptotics are different for vanishing and non-vanishing spatial momentum \vec{k} . We therefore consider both cases in turn.

Asymptotics for $\vec{k} = 0$

In the case of vanishing spatial momentum, we make a power-law ansatz for the second-order differential equation of the chiral components. The asymptotic expansion near the boundary then comprises two independent solutions and is given by

$$\phi_{\pm} = r^{\pm M} (1 + r^{-2z} c_{\pm} + \dots) A_{\pm} + r^{\mp M - z} (1 + r^{-2z} d_{\pm} + \dots) B_{\pm} , \quad (\text{B.7})$$

where we included subleading corrections in the two separate branches. Here, A_{\pm} and B_{\pm} are spinors in the Clifford algebra, and the coefficients c_{\pm} and d_{\pm} are matrices therein. We also observe that (B.7) agrees with the AdS case [36] for $z = 1$, and we note that a subleading term in the A (or B) branch can be more dominant over the leading term in the B (or A) branch for certain ranges of M and z . Employing the expansion (B.7) in the first order Dirac equation (B.5), we derive a local relation between B_{\mp} and A_{\pm} of the form

$$B_{\pm} = -i \frac{\omega \gamma^0}{2M \pm z} A_{\mp} , \quad (\text{B.8})$$

which fixes half of the integration constants. The coefficients of the subleading terms are found as

$$c_{\pm} = \frac{\omega^2}{2z(\pm 2M - z)} , \quad d_{\pm} = \frac{\omega^2}{2z(\mp 2M - 3z)} . \quad (\text{B.9})$$

However, here we need to assume that the subleading terms in the expansion (B.7) are still leading over the subleading terms in the Dirac equation (B.5) which appear in the function V . A careful analysis translates this assumption into the condition

$$z < d - 1 . \quad (\text{B.10})$$

As we can see from (B.8) and (B.9), the above analysis breaks down for $M = \pm z/2$. Moreover, in that case the powers in the asymptotic expansion (B.7) can become equal and so the spinor coefficients cannot be disentangled. The way to treat the case $M =$

$\pm z/2$ is to allow for logarithmic terms in the asymptotic expansion, which lift the degeneracy in the two branches A and B . In particular, for $M = -z/2$ we have

$$\begin{aligned}\phi_+ &= r^{-\frac{z}{2}} A_+ + r^{-\frac{z}{2}} \log(r) B_+ + \dots, \\ \phi_- &= r^{+\frac{z}{2}} \left(1 + c_- r^{-2z} \log(r) + \dots\right) A_- + r^{-3\frac{z}{2}} \left(1 + \dots\right) B_-, \end{aligned} \quad (\text{B.11})$$

with the coefficients of the form

$$B_+ = i\omega\gamma^0 A_- , \quad B_- = \frac{i\omega\gamma^0}{2z} A_+ , \quad c_- = \frac{\omega^2}{2z} . \quad (\text{B.12})$$

For $M = +z/2$, we find

$$\begin{aligned}\phi_+ &= r^{+\frac{z}{2}} A_+ + r^{-3\frac{z}{2}} B_+ + \dots, \\ \phi_- &= r^{-\frac{z}{2}} A_- + r^{-\frac{z}{2}} \log(r) B_- + \dots, \end{aligned} \quad (\text{B.13})$$

together with

$$B_- = -i\omega\gamma^0 A_+ , \quad B_+ = -\frac{i\omega\gamma^0}{2z} A_- . \quad (\text{B.14})$$

Asymptotics for $\vec{k} \neq 0$

For non-zero spatial momenta \vec{k} and generic values of the Dirac mass M , the asymptotic solution to the second-order differential equation (B.6) reads

$$\phi_{\pm} = r^{\pm M} \left(1 + r^{-2} c_{\pm} + \dots\right) A_{\pm} + r^{\mp M-1} \left(1 + r^{-\delta_{\pm}} d_{\pm} + \dots\right) B_{\pm} . \quad (\text{B.15})$$

Note the important difference compared to (B.7), namely that the exponent of the leading order term does not depend on z . The relation between A_{\pm} and B_{\mp} can again be derived via the first-order Dirac equation (B.5). Together with the subleading coefficient c_{\pm} , their form reads

$$B_{\pm} = i \frac{\vec{k} \cdot \vec{\gamma}}{2M \pm 1} A_{\mp} , \quad c_{\pm} = -\frac{\vec{k}^2}{2(\pm 2M - 1)} . \quad (\text{B.16})$$

The subleading behavior in the B-branch turns out to be more complicated. Both the exponent δ_{\pm} and the coefficient d_{\pm} depend on the value of z . We have to distinguish between three separate cases

$$\begin{aligned} 1 < z < 3 : \quad & \delta_{\pm} = z - 1 , \quad d_{\pm} = \frac{\omega}{\vec{k}^2} \left(\frac{\pm 2M + 1}{\pm 2M + z} \right) \vec{\gamma} \cdot \vec{k} , \\ z = 3 : \quad & \delta_{\pm} = 2 , \quad d_{\pm} = \frac{1}{3 \pm 2M} \left(\frac{\vec{k}^2}{2} + \frac{\omega}{\vec{k}^2} (2M \pm 1) \vec{\gamma} \cdot \vec{k} \right) , \\ 3 < z : \quad & \delta_{\pm} = 2 , \quad d_{\pm} = \frac{\vec{k}^2}{2(3 \pm 2M)} . \end{aligned} \quad (\text{B.17})$$

As is clear from (B.16), the case of $M = \pm 1/2$ is again special and needs to be treated separately. Along similar lines as for vanishing momenta, logarithmic terms will appear in the expansion.

B.3 Hydrodynamic limit

We can determine the Green's function analytically in the strict hydrodynamic limit $\omega \rightarrow 0$, $\vec{k} \rightarrow 0$ for arbitrary temperatures T in the region $|M| < z/2$. However, one has to be careful in defining this limit because the $r \rightarrow r_h$ and $\omega \rightarrow 0$ limits do not commute in the differential equation (2.15). Here, we define the limiting procedure as first setting $\vec{k} = 0$, and then taking $\omega \rightarrow 0$ while keeping T finite. Thus, we should use the same boundary condition $\xi(r_h) = i$ as before. The differential equation (2.15) for $\xi(r)$ then becomes

$$r^2 V \partial_r \xi_{\pm} + 2Mr \xi_{\pm} = 0, \quad (\text{B.18})$$

whose solution with the boundary condition $\xi_+(r_h) = i$ reads

$$\xi_{\pm}(r) = i e^{-2M \int_{r_h}^r \frac{dt}{tV(t)}}. \quad (\text{B.19})$$

Substituting then the expression for $V(r)$ given in (2.1) and performing the integral, we find

$$\xi_{\pm}(r) = i \left(\frac{r}{r_h} \right)^{-2M} [1 + V(r)]^{-\frac{4M}{d+z-1}}, \quad (\text{B.20})$$

where we have left the dimension d and the dynamical exponent z unspecified. The self-energy appearing in the Green's function is given by (2.12), and so we obtain

$$G_R(\vec{0}, 0) = \frac{i}{g} 2^{\frac{4M}{d+z-1}} \left(\frac{d+z-1}{4\pi} \right)^{\frac{2M}{z}} T^{-\frac{2M}{z}}, \quad (\text{B.21})$$

where we employed (2.2) to related r_h to the temperature T . Note that this expression is valid for $d+z-1 \neq 0$, which includes the case of interest in this paper, namely $d=4$ and $z=2$.

C Symmetry properties of the Green's function

In this appendix, we discuss symmetries of the self-energy $\Sigma(p)$ defined in (2.12) and of the full Green's function (2.11). This analysis utilizes the properties of the differential equation (2.15) and boundary condition (2.16), which we recall for convenience

$$r^2 V \partial_r \xi_{\pm} + 2Mr \xi_{\pm} = \frac{\omega}{r^{z-1}V} \mp k_3 + \left(\frac{\omega}{r^{z-1}V} \pm k_3 \right) \xi_{\pm}^2, \quad \xi_{\pm}(r_h) = i. \quad (\text{C.1})$$

C.1 Scale-invariant variables

We start our discussion with the scaling properties of (C.1). To simplify our notation, we define the constant $\kappa = \frac{d+z-1}{4\pi}$, and we can distinguish three cases which will be considered in turn.

- We define new variables in the following way

$$\bar{\omega} = \frac{\omega}{k_3^z}, \quad \bar{T} = \frac{T}{k_3^z}, \quad x = \frac{r}{k_3}, \quad (\text{C.2})$$

which leads to

$$V = \sqrt{1 - \left(\frac{x_h}{x}\right)^{d+z-1}}, \quad x_h = \left(\frac{\bar{T}}{\kappa}\right)^{\frac{1}{z}}. \quad (\text{C.3})$$

The differential equations (C.1) then takes the form

$$x^2 V \partial_x \xi_{\pm} + 2Mx \xi_{\pm} = \frac{\bar{\omega}}{x^{z-1}V} \mp 1 + \left(\frac{\bar{\omega}}{x^{z-1}V} \pm 1\right) \xi_{\pm}^2, \quad \xi_{\pm}(x_h) = i, \quad (\text{C.4})$$

and we obtain

$$\xi_{\pm}(\bar{x}_0, \bar{\omega}, \bar{T}) = k_3^{-2M} \xi_{\pm}(r_0, \omega, k_3, T). \quad (\text{C.5})$$

- Along similar lines, we can define another set of variables as follows

$$\bar{k} = \frac{k_3}{\omega^{\frac{1}{z}}}, \quad \bar{T} = \frac{T}{\omega}, \quad y = \frac{r}{\omega^{\frac{1}{z}}}, \quad (\text{C.6})$$

leading to

$$V = \sqrt{1 - \left(\frac{y_h}{y}\right)^{d+z-1}}, \quad y_h = \left(\frac{\bar{T}}{\kappa}\right)^{\frac{1}{z}}, \quad (\text{C.7})$$

as well as

$$y^2 V \partial_y \xi_{\pm} + 2My \xi_{\pm} = \frac{1}{y^{z-1}V} \mp \bar{k} + \left(\frac{1}{y^{z-1}V} \pm \bar{k}\right) \xi_{\pm}^2, \quad \xi_{\pm}(y_h) = i. \quad (\text{C.8})$$

We then find

$$\xi_{\pm}(\bar{y}_0, \bar{k}, \bar{T}) = \omega^{-\frac{2M}{z}} \xi_{\pm}(r_0, \omega, k_3, T). \quad (\text{C.9})$$

- Finally, we can define variables as

$$\bar{\omega} = \frac{\kappa}{T} \omega, \quad \bar{k} = \left(\frac{\kappa}{T}\right)^{\frac{1}{z}} k_3, \quad z = \frac{r}{r_h}, \quad (\text{C.10})$$

where we recall that $r_h = (T/\kappa)^{\frac{1}{z}}$. The above choice implies

$$V = \sqrt{1 - \left(\frac{z_h}{z}\right)^{d+z-1}}, \quad z_h = 1, \quad (\text{C.11})$$

and the differential equations become

$$z^2 V \partial_z \xi_{\pm} + 2Mz \xi_{\pm} = \frac{\bar{\omega}}{r^{z-1}V} \mp \bar{k} + \left(\frac{\bar{\omega}}{r^{z-1}V} \pm \bar{k}\right) \xi_{\pm}^2, \quad \xi_{\pm}(z_h) = i. \quad (\text{C.12})$$

We then have

$$\xi_{\pm}(\bar{z}_0, \bar{\omega}, \bar{k}) = \left(\frac{T}{\kappa}\right)^{-\frac{2M}{z}} \xi_{\pm}(r_0, \omega, k_3, T). \quad (\text{C.13})$$

C.2 Chirality

From equation (C.1) we observe that the spin-up and spin-down components ξ_+ and ξ_- are related by a reflection of the momentum

$$\xi_{\pm}(r, k_3, \omega) = \xi_{\mp}(r, -k_3, \omega) . \quad (\text{C.14})$$

Furthermore, equations (2.12) and (2.11) which relate ξ to the self-energy $\Sigma(p)$ and to the Green's function Σ to G_R do not spoil this symmetry. We thus obtain the following relation for the components of the full Green's function

$$G_R^{\pm}(\vec{k}, \omega) = G_R^{\mp}(-\vec{k}, \omega) . \quad (\text{C.15})$$

This symmetry can easily be understood in the relativistic case of $z = 1$: since ψ_{\pm} correspond to left- and right-handed spinors in the boundary theory, the helicity h given by $h = \vec{\sigma} \cdot \vec{k} / |\vec{\sigma} \cdot \vec{k}|$ is conserved. In the non-relativistic case $z \neq 1$, this may be viewed as a generalization of chirality.

C.3 Particle-hole symmetry

From equation (C.1) we also observe that the components ξ_+ and ξ_- enjoy the additional symmetry

$$\xi_{\pm}(r, k_3, \omega) = -\xi_{\pm}^*(r, -k_3, -\omega) , \quad (\text{C.16})$$

where ξ^* denotes the complex conjugate of ξ . Noting then again that (2.12) and (2.11) preserve this property as well, we see that also the full Green's function satisfies

$$G_R^{\pm}(\vec{k}, \omega) = -\left(G_R^{\pm}(-\vec{k}, -\omega)\right)^* = -\left(G_R^{\mp}(\vec{k}, -\omega)\right)^* , \quad (\text{C.17})$$

where we used the chirality symmetry in the last equality. This symmetry can be seen as a *particle-hole* symmetry since it relates components with frequency ω to components with frequency $-\omega$.

C.4 A symmetry relating M to $-M$

In the case when the operator under consideration is right handed, that is \mathcal{O}_+ , the source is identified with the boundary value of ψ_- instead. Hence, the transfer matrix $\xi(r, p)$ defined in (2.7) is replaced with

$$\zeta_+ = i \frac{\psi_{+,1}}{\psi_{-,1}} , \quad \zeta_- = i \frac{\psi_{+,2}}{\psi_{-,2}} , \quad (\text{C.18})$$

where $\psi_{1,2}$ are the up and down components of the spinors defined in (2.5) and (2.6). Let us now consider ζ_+ for definiteness, the discussion for ζ_- is analogous. First, we see that $\zeta_+ = -1/\xi_+$, and we easily verify that the equation for ζ_+ is exactly the same

equation (2.15) that ξ_+ satisfies, but with the replacement $k_3 \rightarrow -k_3$ and $M \rightarrow -M$. In addition, the boundary conditions (2.16) for ζ_+ and ξ_+ are the same, and therefore we have the following symmetry property:

$$\zeta_+(r, M, \omega, k_3) = -\xi_+^{-1}(r, M, \omega, k_3) = \xi_+(r, -M, \omega, -k_3) . \quad (\text{C.19})$$

On the other hand, equation (2.12) is now replaced with

$$\Sigma(p) = -g \lim_{r_0 \rightarrow \infty} r_0^{-2M} \zeta(r_0, p) , \quad (\text{C.20})$$

and so we arrive at

$$\Sigma(-M, -\vec{k}, \omega) = -g^2 \Sigma^{-1}(+M, +\vec{k}, \omega) . \quad (\text{C.21})$$

However, note that strictly speaking this derivation is valid only within the range $-z/2 < M < z/2$ for $\vec{k} = 0$ and within $-\frac{1}{2} < M < \frac{1}{2}$ for $\vec{k} \neq 0$. Outside this region one needs to subtract divergent terms to obtain the Green's functions from ξ and ζ , which can be shown by observing that the counter terms that are needed in order to renormalize G_R also respect this symmetry.

As a consequence of this symmetry, conditions on the functions $s_{1,M}$ and $s_{2,M}$ from (3.15) can be obtained as

$$\begin{aligned} s_{1,-M}(u) s_{2,+M}(u) &= s_{1,+M}(u) s_{2,-M}(u) , \\ s_{1,+M}(u) s_{1,-M}(u) - s_{2,+M}(u) s_{2,-M}(u) &= -1 . \end{aligned} \quad (\text{C.22})$$

Solving (C.22) in favor of $s_{2,M}(u)$, we find for non-degenerate values of $s_{1,M}$ and $s_{2,M}$ that

$$s_{2,M}(u) = s_{1,M}(u) \sqrt{1 + \frac{1}{s_{1,M}(u) s_{1,-M}(u)}} , \quad (\text{C.23})$$

and by using (C.23) in (3.15) we arrive at

$$G_R(\vec{k}, \omega) = - \frac{1}{\omega - \frac{1}{\lambda} \vec{\sigma} \cdot \vec{k} k^{z-1} - g k^{2M} s_{1,M}(u) \left(1 + \frac{\vec{\sigma} \cdot \vec{k}}{k} \sqrt{1 + \frac{1}{s_{1,M}(u) s_{1,-M}(u)}} \right)} . \quad (\text{C.24})$$

Thus, we have reduced the problem to the determination of a single function $s_{1,M}(u)$ of a single variable $u = \omega/k^z$ for two values of M .

C.5 A symmetry of the momentum distribution

Using the above symmetries of the Green's function in combination with the sum rule, we can easily derive a convenient identity for the momentum distribution (4.1). Namely,

we have

$$\begin{aligned}
N_{\vec{k}}^{\pm} + N_{-\vec{k}}^{\pm} &= \frac{1}{\pi} \int_{-\infty}^{+\infty} d\omega \left[\text{Im } G_R^{\pm}(\vec{k}, \omega) n_F(\omega) + \text{Im } G_R^{\pm}(-\vec{k}, \omega) n_F(\omega) \right] \\
&= \frac{1}{\pi} \int_{-\infty}^{+\infty} d\omega \text{Im } G_R^{\pm}(\vec{k}, \omega) [n_F(\omega) + n_F(-\omega)] \\
&= 1,
\end{aligned} \tag{C.25}$$

where we used (C.17) in the second line, and the sum rule (2.18) and $n_F(\omega) + n_F(-\omega) = 1$ in the third line.

D Pole structure and sum rules

In this appendix, we analyze the pole structure of the Green's function, and we show analytically for the case of vanishing temperature that the sum rule is satisfied. The situation of vanishing momenta can be analyzed for all values of z , however, analytic expressions for the Green's function for $k \neq 0$ are only available for $z = 1$. We therefore restrict the latter computations to the relativistic case.

D.1 Pole structure

We now consider the pole structure of the Green's function in the case of vanishing momentum and zero temperature. Employing (3.13), we see that there is a trivial pole at $\omega = 0$ which corresponds to the free particle. A non-trivial pole is determined by the equation

$$0 = 1 - g c_z \omega^{\frac{2M}{z}-1} e^{-i\pi(\frac{M}{z} + \frac{1}{2})}. \tag{D.1}$$

Using then polar coordinates $\omega = |\omega|e^{i\theta}$, we find for the imaginary part of (D.1) the equation

$$0 = \sin \left[\frac{2M-z}{z} \theta - \frac{2M+z}{z} \frac{\pi}{2} \right], \tag{D.2}$$

with solutions

$$\theta_n = \frac{\pi}{2} - \frac{n+1}{z-2M} \pi z, \quad n \in \mathbb{Z}. \tag{D.3}$$

If we want to satisfy the Kramers-Kronig relations, that is the requirement of no poles in the upper half-plane, we have to make sure that there is no solution to (D.3) in the range $(0, +\pi)$ for all values of n , that is

$$\theta_n \notin (0, +\pi) \quad \forall n. \tag{D.4}$$

By noting that $0 < z - 2M < 2z$ for $|M| < \frac{z}{2}$, we see that the only possible solution in the forbidden range is given by $n = -1$ with $\theta_{-1} = \frac{\pi}{2}$. All the other solutions are

already out of the forbidden range, but yield poles on the non-principal sheets of the complex plane. For example, for $n = -2$ the pole resides in $\theta_{-2} \in (\pi, +\infty)$ depending on the value of M , while for $n = 0$ it is in the range $\theta_0 \in (0, -\infty)$.

Let us now turn to the real part of (D.1). Employing again polar coordinates and using (D.3), we arrive at the equation

$$0 = 1 - g c_z |\omega|^{\frac{2M}{z}-1} (-1)^n, \quad (\text{D.5})$$

which is solved by

$$|\omega| = |g c_z|^{\frac{z}{z-2M}}, \quad \text{sign}(g c_z) = (-1)^n. \quad (\text{D.6})$$

It is now clear that in order to avoid the pole $\theta_{-1} = \frac{\pi}{2}$ on the principal sheet, we have to require

$$g c_z > 0. \quad (\text{D.7})$$

Furthermore, all other possible poles of the Green's function (specified by (D.3) and (D.6)) with even n are either on the lower-half of the principal sheet or on non-principal sheets of the complex plane. Therefore, for the choice of sign (D.7) the Kramers-Kronig relations are satisfied.

For the case of vanishing temperature and $z = 1$, analytic results also for $\vec{k} \neq 0$ are available. As it has been explained in [14], the analytic continuation of the Green's function for non-vanishing momenta is given by replacing $\omega \rightarrow \sqrt{\omega^2 - k^2}$. This implies that there are poles at $\omega = \pm |\vec{k}|$, with the pole in the lower half-plane corresponding to the retarded Green's function. Furthermore, for the choice of sign (D.7), there are no additional poles in the upper half-plane of the principal sheet. Therefore, we have shown that in the relativistic case the Kramers-Kronig relations are satisfied also for non-vanishing momenta.

D.2 Sum rule for $\vec{k} = 0$

The Green's function in the case of vanishing temperature and momenta can be found in equation (3.13). The spectral density (2.17) then reads

$$\rho(\vec{k} = 0, \omega) = \frac{1}{2\pi} \text{Im Tr}[G_R(\vec{k} = 0, \omega)] = -\frac{1}{\pi} \text{Im} \left[\frac{1}{\omega - g c_z \omega^{\frac{2M}{z}} e^{-i\pi(\frac{M}{z} + \frac{1}{2})}} \right]. \quad (\text{D.8})$$

Introducing the following constants

$$\mathcal{B} = g c_z \cos \left[\pi \left(\frac{M}{z} + \frac{1}{2} \right) \right], \quad \mathcal{C} = g c_z \sin \left[\pi \left(\frac{M}{z} + \frac{1}{2} \right) \right], \quad (\text{D.9})$$

we express the spectral density as

$$\rho(\vec{k} = 0, \omega) = +\frac{1}{\pi} \frac{\mathcal{C} \omega^{\frac{2M}{z}}}{\left(\omega - \mathcal{B} \omega^{\frac{2M}{z}} \right)^2 + \left(\mathcal{C} \omega^{\frac{2M}{z}} \right)^2}. \quad (\text{D.10})$$

Next, to compute the sum rule we recall the symmetry $\rho(\vec{k}, -\omega) = \rho(\vec{k}, +\omega)$. We then have

$$\begin{aligned} \int_{-\infty}^{+\infty} d\omega \rho(\vec{k} = 0, \omega) &= 2 \int_0^{+\infty} d\omega \rho(\vec{k} = 0, \omega) \\ &= \frac{2}{\pi} \frac{z}{2M - z} \arctan \left[-\frac{\mathcal{B}}{\mathcal{C}} + \omega^{\frac{2M}{z}} \frac{\mathcal{B}^2 + \mathcal{C}^2}{\mathcal{C}} \right]_{\omega=0}^{\omega=\infty} \\ &= \frac{2M}{2M - z} - \frac{z}{2M - z} \text{sign}(gc_z) . \end{aligned} \quad (\text{D.11})$$

Therefore, for the choice of sign (3.3) we indeed find that the sum rule (2.18) for vanishing momenta and vanishing temperature is satisfied

$$\int_{-\infty}^{+\infty} d\omega \rho(\vec{k} = 0, \omega) = 1 . \quad (\text{D.12})$$

D.3 Sum rule for $\vec{k} \neq 0$

In the case of $z = 1$ and vanishing temperature, the expression for the Green's function was given in equation (3.1) together with (3.4). To compute the sum rule in this situation, we write

$$\int_{-\infty}^{+\infty} d\omega \rho(\vec{k}, \omega) = \frac{2}{2\pi} \int_0^{+\infty} d\omega \text{Im} \frac{-2\omega}{p^2 \left(1 - gc_1 e^{-i\pi(M+\frac{1}{2})} p^{2M-1} \right)} , \quad (\text{D.13})$$

with p as in (3.4). Next, we take (3.4) into account to obtain

$$\begin{aligned} \int_{-\infty}^{+\infty} d\omega \rho(\vec{k}, \omega) &= + \frac{1}{\pi} \int_0^{+|\vec{k}|} d\omega \text{Im} \frac{-2\omega}{(\omega^2 - k^2) (1 + gc_1 (k^2 - \omega^2)^{2M-1})} \\ &\quad + \frac{1}{\pi} \int_{+|\vec{k}|}^{+\infty} d\omega \text{Im} \frac{-2\omega}{(\omega^2 - k^2)^2 \left(1 - gc_1 e^{-i\pi(M+\frac{1}{2})} (\omega^2 - k^2)^{2M-1} \right)} , \end{aligned} \quad (\text{D.14})$$

where we note that the first term on the right-hand side vanishes due to taking the imaginary part. Performing then the change of variables $x = \omega^2 - k^2$, we find

$$\int_{-\infty}^{+\infty} d\omega \rho(\vec{k}, \omega) = -\frac{1}{\pi} \int_0^\infty dx \text{Im} \left[\frac{1}{x^2 - gc_1 x^{2M+1} e^{-i\pi(M+\frac{1}{2})}} \right] . \quad (\text{D.15})$$

After a further change of variables $y = x^2$, we arrive at

$$\int_{-\infty}^{+\infty} d\omega \rho(\vec{k}, \omega) = -\frac{2}{\pi} \int_0^\infty dy \text{Im} \left[\frac{1}{y - gc_1 y^{2M} e^{-i\pi(M+\frac{1}{2})}} \right] , \quad (\text{D.16})$$

which is a special case of the computation in appendix D.2. We therefore find

$$\int_{-\infty}^{+\infty} d\omega \rho(\vec{k}, \omega) = 1 . \quad (\text{D.17})$$

E WKB calculation

In this appendix, we present details of the calculation leading to equation (4.22). We consider the case $M > 0$, $T = 0$ and without loss of generality make the assumptions $\omega > 0$, $\vec{k} = (0, 0, k_3)$ and $k_3 > 0$. Furthermore, we consider the spin-up component of the Green's function, the other cases can be worked out using the symmetries outlined in appendix C.

The parameter which we take large in the WKB approximation is $\bar{k} \equiv k/\sqrt{\omega}$, and the fluctuation equations of u_{\pm} are given by (B.6) which yield

$$\begin{aligned} 0 &= \partial_r^2 u_{\pm} + \left[\frac{2}{r} + \frac{\omega}{r^2 \tilde{k}^2} \left(-\frac{\omega}{r} \pm k \right) \right] \partial_r u_{\pm} + \left[-\frac{\tilde{k}^2}{r^4} - \frac{M\omega}{r^3 \tilde{k}^2} \left(\mp \frac{\omega}{r} + k \right) - \frac{M(M \pm 1)}{r^2} \right] u_{\pm} \\ &\equiv \partial_r^2 u_{\pm} + f_1^{\pm}(r) \partial_r u_{\pm} + f_2^{\pm}(r) u_{\pm} , \end{aligned} \quad (\text{E.1})$$

where \tilde{k} had been defined in (B.2). This expression can be put in the Schrödinger form by the transformation

$$u_{\pm} = h_{\pm}(r) y_{\pm} , \quad h_{\pm}(r) = \exp \left(-\frac{1}{2} \int^r dt f_1^{\pm}(t) \right) . \quad (\text{E.2})$$

A suitable parametrization of the extra spatial coordinate in the limit $\bar{k} \gg 1$ is given in terms of $x = r/\sqrt{\omega}$. In this variable, the transformation (E.2) reads as follows

$$h_{\pm}(x) = \frac{\sqrt{x\bar{k} \pm 1}}{x^{3/2}} . \quad (\text{E.3})$$

Then, the Schrödinger equations become

$$\frac{1}{\bar{k}^2} \frac{d^2}{dx^2} y_{\pm}(x) - V_{\pm}(x) y_{\pm}(x) = 0 , \quad (\text{E.4})$$

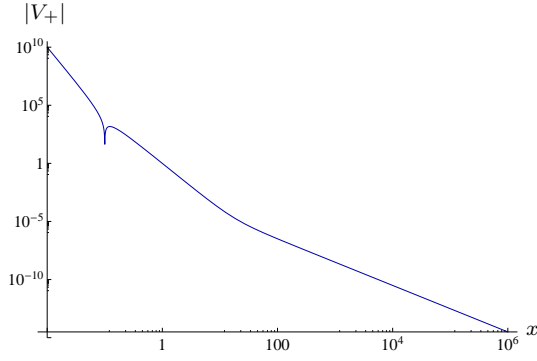
with the effective Schrödinger potentials

$$V_{\pm}(x) = \frac{1}{x^4} - \frac{1}{\bar{k}^2 x^6} + \frac{(\frac{1}{2} \pm M)(\frac{3}{2} \pm M)}{\bar{k}^2 x^2} + \frac{3}{4} \frac{1}{(x\bar{k} \pm 1)^2} - \frac{\frac{3}{2} \pm M}{x\bar{k}(x\bar{k} \pm 1)} . \quad (\text{E.5})$$

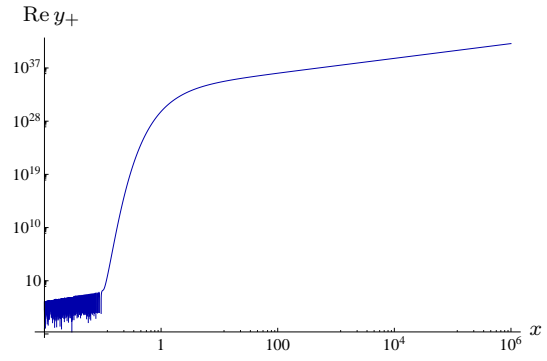
In addition to solving the above Schrödinger equation for y_+ analytically using the WKB approximation as described below, we have also solved it numerically. The numerical results are presented in figure 18, which we discuss as we proceed in our computation.

We now focus on the plus component of the fluctuations. In the region close to the horizon at $x \ll 1$, and in the limit $\bar{k} \gg 1$, the relevant piece of the potential for y_+ reduces to

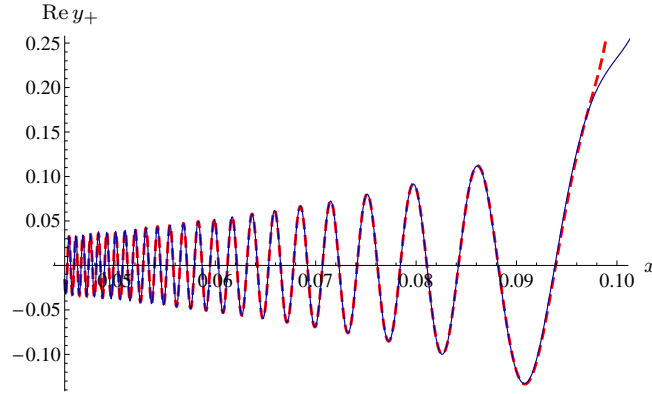
$$V_+ \approx \frac{1}{x^4} - \frac{1}{\bar{k}^2 x^6} , \quad (\text{E.6})$$



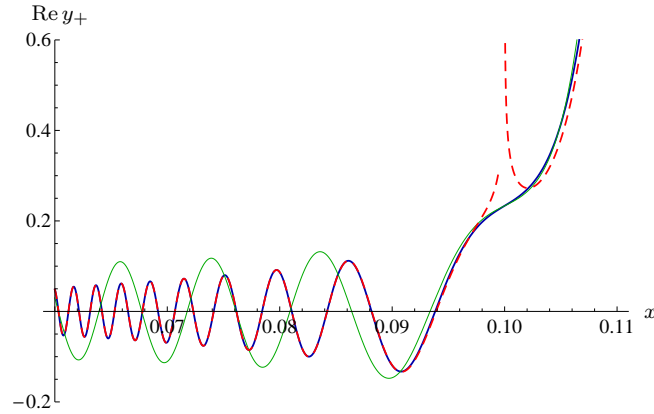
(a) Double-logarithmic plot of $|V_+|$. The turning point is approximately $x_{\text{tp}} = 1/10$.



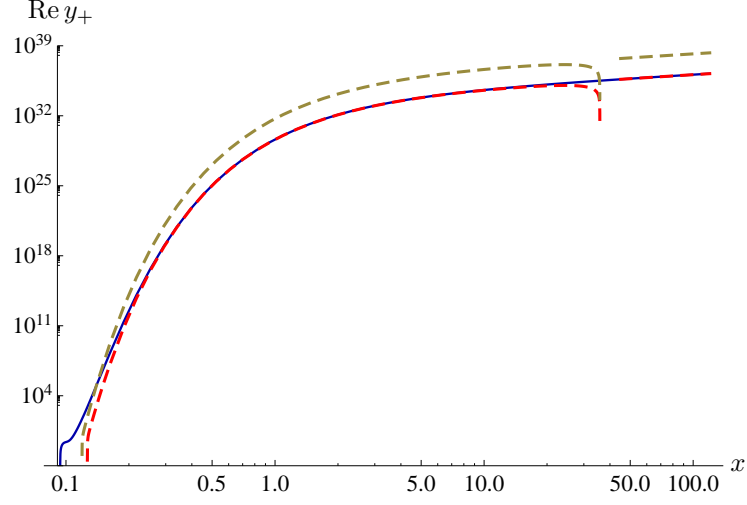
(b) Double-logarithmic plot of $\text{Re } y_+$. The imaginary part is similar but with a minus sign. Left of the turning point there is oscillatory behavior, the other regions exhibit exponential behavior.



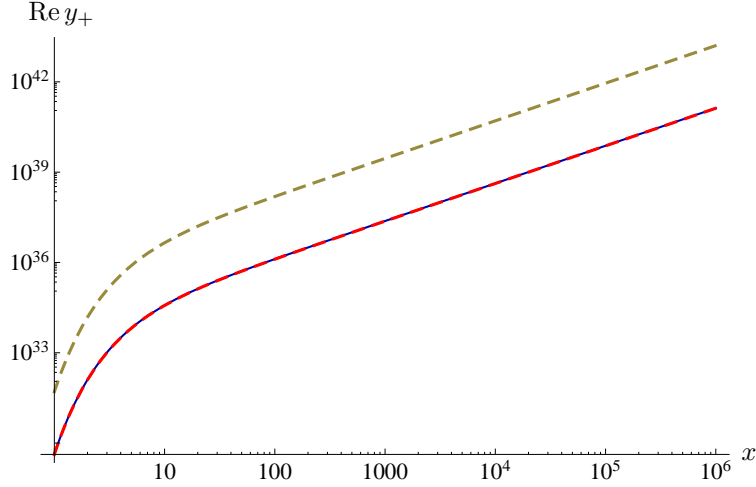
(c) Near-horizon oscillatory behavior. Until the turning point, there is good agreement between the numerical (blue) and WKB (red) result. After that, the WKB solution is no longer valid.



(d) Matching procedure near the classical turning point. The dashed red curves are the WKB results, the solid blue curve is the numerical solution, and the solid green curve is the Airy function solution needed in the matching formulae which agrees with the numerics around the turning point.



(e) Double-logarithmic plot of $\text{Re } y_+$ in the classically disallowed region. The solid blue curve is the numerical solution, the dashed lines are the analytic solution (E.12) for exponents $\alpha = \pi/2$ (red) and $\alpha = 5/3$ (yellow). The former agrees with the numerical result, the latter deviates from it. For large x , the WKB results are not valid anymore but are taken over by the Bessel-function solutions plotted on the far right, which overlap with the WKB solution for smaller x .



(f) Double-logarithmic plot of the near-boundary region, where the Bessel-function solution is valid. The blue curve is the numerical result, the dashed curves are the analytic result (E.16) for exponents $\alpha = \pi/2$ (red) $\alpha = 5/3$ (yellow).

Figure 18: Illustration of the numerical WKB calculation for $\bar{k} = 10$ and $M = 1/4$.

from which we see that there is a turning point located at approximately $x_{\text{tp}} = 1/\bar{k}$. In the region $x < x_{\text{tp}}$ the WKB wavefunction is given by

$$y_+(x) = \frac{A_1}{(-V_+)^{1/4}} e^{+i\bar{k} \int_{x_{\text{tp}}}^x dt \sqrt{-V_+(t)}} + \frac{A_2}{(-V_+)^{1/4}} e^{-i\bar{k} \int_{x_{\text{tp}}}^x dt \sqrt{-V_+(t)}} , \quad (\text{E.7})$$

with V_+ obtained from (E.6). The solution for $x < x_{\text{tp}}$ is compared with the numerics in figure 18(c). This solution should be connected to the solution near the horizon at $x \rightarrow 0$, which yields the condition that $A_2 = 0$. We will not need to fix the coefficient A_1 because, as shown below, it will cancel out in the final answer. The next step is to continue the solution from the region $x < x_{\text{tp}}$ to the region $x > x_{\text{tp}}$ by using the well-known WKB connection formula, see for example [25, 26]. This procedure is illustrated in figure 18(d). One important subtlety here is that in the final answer we will need second-order corrections to the WKB formula as well. Noting that the parameter that plays the role of \hbar here is $1/\bar{k}$, we find,

$$y_+(x) = \frac{A_1 e^{-i\frac{\pi}{4}}}{V_+(x)^{\frac{1}{4}}} \left[e^{+\bar{k} \int_{x_{\text{tp}}}^x dt \sqrt{V_+(t)}} \left(1 + \frac{\sigma_+(x)}{\bar{k}} \right) + \frac{i}{2} e^{-\bar{k} \int_{x_{\text{tp}}}^x dt \sqrt{V_+(t)}} \left(1 + \frac{\sigma_-(x)}{\bar{k}} \right) \right] , \quad (\text{E.8})$$

where the second-order WKB coefficients can be found as (see for example [37])

$$\sigma_{\pm} = \pm \int_{x_{\text{tp}}}^x dt \left(\frac{1}{8} \frac{V_+''(t)}{V_+(t)^{\frac{3}{2}}} - \frac{5}{32} \frac{V_+'(t)^2}{V_+(t)^{\frac{5}{2}}} \right) . \quad (\text{E.9})$$

This is our WKB solution in the classically disallowed region $x_{\text{tp}} \ll x < \infty$, where we can now approximate the potential by⁶

$$V_+ \rightarrow \frac{1}{x^4} + \frac{M(M+1)}{\bar{k}^2 x^2} + \frac{M}{x^3 \bar{k}^3} - \frac{1}{x^6 \bar{k}^2} . \quad (\text{E.10})$$

To the desired order in \bar{k} , (E.9) then yields

$$\pm \sigma_{\pm} = \frac{1}{4\bar{k}^2 x} + \frac{M(M+1)x^3}{4\bar{k}^2} + \frac{Mx^2}{8\bar{k}^3} - \frac{1}{4\bar{k}} + \dots , \quad (\text{E.11})$$

and with this approximation we find from (E.8) the following solution, which is illustrated in figure 18(e),

$$y_+(x) = A_1 x e^{-i\frac{\pi}{4}} \left(e^{-\frac{\bar{k}}{x}} e^{\bar{k}^2} \kappa_+(x) + \frac{i}{2} e^{+\frac{\bar{k}}{x}} e^{-\bar{k}^2} \kappa_-(x) \right) , \quad (\text{E.12})$$

where we employed

$$x_{\text{tp}} = \frac{1}{\bar{k}} - \frac{M(M+1)}{\bar{k}^5} + \dots , \quad (\text{E.13})$$

⁶Note that we keep the leading $1/\bar{k}^2$ corrections in order to be consistent with keeping the σ_{\pm} corrections in (E.8).

and have defined the functions

$$\kappa_{\pm} \approx \left[1 \pm \frac{M(M+1)x}{2\bar{k}} \pm \frac{M}{2\bar{k}^2} \log(\bar{k}x) \right] \left[1 - \frac{M(M+1)x^2}{4\bar{k}^2} + \frac{1}{4x^2\bar{k}^2} \right] \left[1 \mp \frac{1}{4\bar{k}^2} \pm \frac{1}{4\bar{k}^3x} \right]. \quad (\text{E.14})$$

Note that here we only show terms up to the desired $1/\bar{k}^2$ order. The terms in the first brackets come from the corrections to the exponential term in (E.8), the second terms come from the expansion of the $V_+^{-1/4}$ and the last ones from the σ_{\pm} corrections in (E.8).

The next step is to connect this solution to the region near the boundary at $x \rightarrow \infty$ and read off the Green's function from the coefficients of the near-boundary expansion. Note that the WKB approximation in the near-boundary limit $x \rightarrow \infty$ fails, because in this region we have an inverse-square potential which does not satisfy the WKB condition $|V'_+(x)/V_+^{\frac{3}{2}}(x)| \ll 1$. Thus, one has to solve the Schrödinger equation with the potential (E.10) exactly and connect it to the solution (E.12) in the overlapping region $\bar{k}/x \gg 1$. In the latter, we can approximate the potential as

$$V_+(x) \approx \frac{1}{x^4} + \frac{M(M+1)}{x^2\bar{k}^2}, \quad (\text{E.15})$$

and the exact solution to the Schrödinger equation with this potential is found in terms of the Bessel functions as

$$y_+(x) = \sqrt{x} \left[B_1 I_{-\frac{1}{2}-M} \left(\frac{\bar{k}}{x} \right) + B_2 I_{\frac{1}{2}+M} \left(\frac{\bar{k}}{x} \right) \right], \quad (\text{E.16})$$

where $B_{1,2}$ are coefficients so far undetermined. This solution is also shown in figure 18(e). In order to connect it to (E.12), we use the well-known asymptotic formula for the Bessel functions for large values of the variable \bar{k}/x and we find

$$y_+(\bar{k}/x) \rightarrow \frac{x}{\sqrt{2\pi\bar{k}}} \left[B_1 \left(e^{\frac{\bar{k}}{x}} + e^{-i\pi M} e^{-\frac{\bar{k}}{x}} \right) + B_2 \left(e^{\frac{\bar{k}}{x}} - e^{i\pi M} e^{-\frac{\bar{k}}{x}} \right) \right] + \dots, \quad (\text{E.17})$$

from which we obtain

$$B_1 = \frac{A_1 e^{-i\frac{\pi}{4}} \sqrt{2\pi\bar{k}}}{2 \cos(\pi M)} \left(\frac{i}{2} e^{i\pi M} e^{-\bar{k}^2} \epsilon_- + e^{\bar{k}^2} \epsilon_+ \right). \quad (\text{E.18})$$

We shall not need the expression for B_2 , and the constants ϵ_{\pm} are defined as follows

$$\epsilon_{\pm} = 1 \pm \frac{M}{2\bar{k}^2} \log \bar{k} \mp \frac{1}{4\bar{k}^2}. \quad (\text{E.19})$$

On the other hand, using equation (E.3) together with the series expansion of the Bessel functions near $x \rightarrow \infty$ given in (E.16), we find that

$$u_+(x) \approx \bar{A}_+^u x^M + \dots, \quad (\text{E.20})$$

where the superscript u denotes that we consider the spin-up component of the spinor \bar{A}_+ . This is in accord with the asymptotic solution given in appendix B.2 with

$$\bar{A}_+^u = B_1 \frac{\bar{k}^{-M} 2^{M+\frac{1}{2}}}{\Gamma(\frac{1}{2} - M)}, \quad (\text{E.21})$$

where B_1 is determined by (E.18).

The next step in the calculation of the Green's function is to obtain the WKB solution for the u_- component in (E.1). One way would be to apply the same steps as above in the WKB calculation, however this process is more difficult than the u_+ case, since the potential for u_- shown in (E.5) has extra poles at $x = 1/\bar{k}$. Fortunately, u_- is related to u_+ by the first order differential equation (B.5), and so we can use the result for u_+ to obtain u_- directly. Equation (B.5), specified to the spin-up component and in the rescaled variables x and \bar{k} , reads

$$u_- = \frac{i}{\bar{k} + \frac{1}{x}} (x^2 \partial_x - Mx) u_+, \quad (\text{E.22})$$

which for the solution (E.8) becomes

$$u_- = \frac{i}{\bar{k} + \frac{1}{x}} \left[x^2 \left(\frac{h'_+(x)}{h_+(x)} - \frac{1}{4} \frac{V'(x)}{V(x)} \right) - Mx \right] u_+ + A(x), \quad (\text{E.23})$$

where the function $A(x)$ is given by

$$A(x) = \frac{ix^2 h_+}{\bar{k} + \frac{1}{x}} \frac{A_1 e^{-i\pi/4}}{V_+^{1/4}(x)} \left[e^{+\bar{k} \int_{x_{\text{tp}}}^x dt \sqrt{V_+(t)}} \left(\bar{k} V_+^{\frac{1}{2}}(x) \left(1 + \frac{\sigma_+(x)}{\bar{k}} \right) + \frac{\sigma'_+(x)}{\bar{k}} \right) - \frac{i}{2} e^{-\bar{k} \int_{x_{\text{tp}}}^x dt \sqrt{V_+(t)}} \left(\bar{k} V_+^{\frac{1}{2}}(x) \left(1 + \frac{\sigma_-(x)}{\bar{k}} \right) - \frac{\sigma'_-(x)}{\bar{k}} \right) \right]. \quad (\text{E.24})$$

On the other hand, very close to the boundary, we can approximate the Schrödinger potential as

$$V_-(x) \approx \frac{1}{x^4} + \frac{M(M-1)}{x^2 \bar{k}^2}, \quad (\text{E.25})$$

and again, using this potential in the Schrödinger-like equation for y_- , the exact solution reads

$$y_-(x) = \sqrt{x} \left[C_1 I_{-\frac{1}{2}+M} \left(\frac{\bar{k}}{x} \right) + C_2 I_{\frac{1}{2}-M} \left(\frac{\bar{k}}{x} \right) \right], \quad (\text{E.26})$$

with $C_{1,2}$ some coefficients to be determined. Similarly as above, expanding the Bessel functions in this solution for large x/\bar{k} , and using $u_- = h_-(x)y_-(x)$ with (E.3), we find

$$u_-(x) \approx \bar{A}_-^u x^{-M} + \dots, \quad (\text{E.27})$$

with

$$\bar{A}_-^u = C_1 \frac{\bar{k}^M 2^{-M+\frac{1}{2}}}{\Gamma(\frac{1}{2} + M)}. \quad (\text{E.28})$$

Next, we employ the asymptotic formula for the Bessel functions in (E.26) for large \bar{k}/x to match to the WKB result (E.23). In this matching, we also need to use (E.24), (E.12) and the relation $u_{\pm}(x) = h_{\pm}(x)y_{\pm}(x)$ with (E.3). The result is

$$C_1 = \frac{A_1 e^{-i\frac{\pi}{4}} \sqrt{2\pi\bar{k}}}{2 \cos(\pi M)} \left(\frac{1}{2} e^{-i\pi M} e^{-\bar{k}^2} \gamma_- + i e^{\bar{k}^2} \gamma_+ \right), \quad (\text{E.29})$$

and we shall not need the expression for C_2 . The constants γ_{\pm} read

$$\gamma_{\pm} = 1 \pm \frac{M}{2\bar{k}^2} \log \bar{k} \mp \frac{3}{4\bar{k}^2}. \quad (\text{E.30})$$

Finally, as can be inferred for instance from (2.7), the spin-up eigenvalue of the matrix ξ is proportional to the ratio of u_- and u_+ . Thus, using its definition from (2.12), the spin-up component of the self-energy is proportional to the ratio of the constants \bar{A}_- and \bar{A}_+ . To be more precise,

$$\Sigma^+ = -i g \omega^M \frac{\bar{A}_-^u}{\bar{A}_+^u} = -i g 2^{-2M} k^{2M} \frac{\Gamma(\frac{1}{2} - M)}{\Gamma(\frac{1}{2} + M)} \frac{C_1}{B_1}. \quad (\text{E.31})$$

If we then use (E.18), (E.21), (E.28) and (E.29), we obtain the final result for the spin-up component of the self-energy as it was given in (4.21).

Before closing this appendix, let us remark the following. As explained at the beginning of section 4.3, the numerical prefactors both of the terms in ϵ_{\pm} and γ_{\pm} and the one in the exponent, i.e., the number α in the end result $\text{Im } \Sigma \sim \exp(-\alpha k^2/\omega)$, are difficult to determine analytically. This is because the integral in (E.8) leading to κ_{\pm} cannot be done analytically. In the calculation presented here, we have approximated the potential V_+ by taking into account only the three most relevant terms in the region $x > x_{\text{tp}}$, which leads to $\alpha = 2$ as in (E.12). This is reasonable away from the turning point, but close to the lower integration limit x_{tp} the x^{-6} term has a considerable contribution not taken into account. Furthermore, the turning point itself is determined by setting to zero an approximation to the effective potential. To do better, we should use the full potential everywhere, and expand the integrand as a series in $1/\bar{k}$. However, this series is not uniformly convergent because the integration limit x_{tp} depends on $1/\bar{k}$. In practice we can terminate the series at some point, which leads to a different approximation of the numerical prefactors. For example, if the full effective potential is taken into account and the integrand $\sqrt{V_+}$ is expanded in $1/\bar{k}$, the next contribution of order \bar{k}^2 is $-\bar{k}^2/6$, which results in $\alpha = 5/3$, and thus in a decay rate of $\Gamma \sim \exp(-5k^2/3\omega)$. Furthermore, we can do the same for the integral in the σ_{\pm} from (E.9). Also, in the prefactor $V_+^{-1/4}$ the full potential can be taken into account. Then, the $\mathcal{O}(1/\bar{k}^2)$ terms in γ_{\pm} and ϵ_{\pm} also

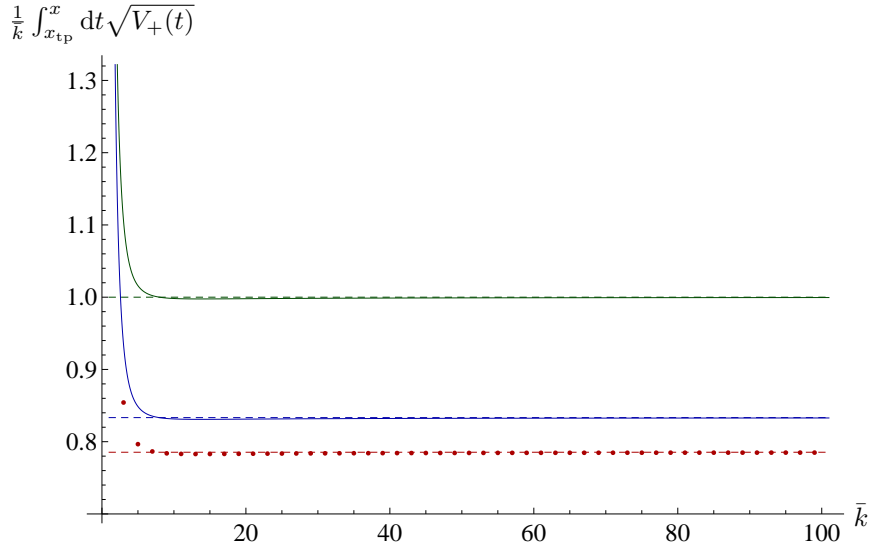


Figure 19: Plot of $\frac{1}{\bar{k}} \int_{x_{\text{tp}}}^x dt \sqrt{V_+(t)}$ as a function of \bar{k} , which appears in the exponents of (E.8) that determine the quasi-particle decay rate. For large \bar{k} , the \bar{k}^2 term in the exponent is dominant and therefore the curves converge to the numerical prefactor of this term. In this way, the correct factor can be determined numerically. Here, we have taken $M = +1/4$ and $x = 20$ as the upper integration limit, which is an intermediate value of x . The value of the integral is independent of M for $\bar{k} \rightarrow \infty$.

The solid green curve is the result when the two most relevant terms in the potential are taken into account, leading to $\alpha = 2$ (dashed green line). The solid blue curve is the resulting expression when the integration over the full potential is done analytically before terminating the series in \bar{k} , leading to the $\alpha = 5/3$ prefactor of the \bar{k}^2 term (the dashed blue line is the number $5/6$). The red dots show the value of the integral when it is performed numerically, and the dashed red line is the number $\pi/4$. The former converge to the latter curve, which is evidence that the value of the integral is approximately $\pi/4\bar{k}^2 + \dots$. Thus, the exponent that determines the quasi-particle decay rate in the final WKB result will be $\exp(-\pi k^2/2\omega)$.

receive extra corrections. Still, we are left with an approximation of the exact numbers, and it is not clear whether there occur cancellations between terms later on in the $1/\bar{k}$ expansion.

The integral over $\sqrt{V_+}$ can however be performed numerically. Figure 19 compares the numerical and analytic values of the number $\alpha/2$ that appears in the exponent. The numerical result is approximately $\pi/4$, whereas our improved approximated analytic calculation yields $5/6$. When we use $\alpha = \pi/2$ in the exponent, the resulting self-energy agrees more accurately with the numerical results, as is shown in figures 15 and 18.

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